

# Characterization of the lonosphere in the Seychelles

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#### Overview

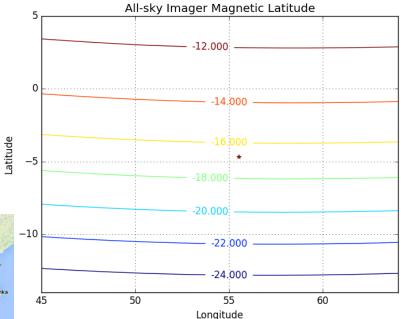
- Location
- Equipment
- Analysis
- Data:
  - Typical Night
  - Collection
  - Monthly
  - Seasonal
  - Depletion
- Events
  - March 17<sup>th</sup>, 2015
  - April 2<sup>nd</sup>, 2016
  - March 30<sup>th</sup>, 2017
- Summary



# The Seychelles

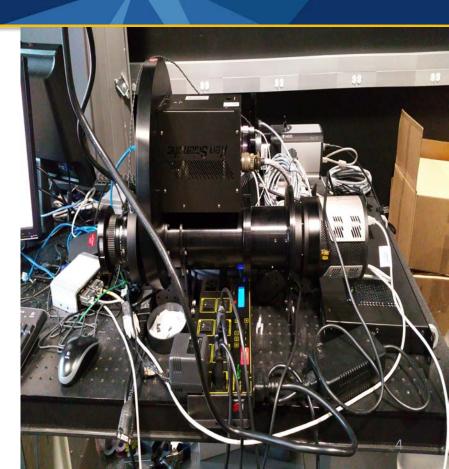
- Island chain in Indian Ocean
- Low Latitude Station
- Southern Hemisphere
- Near southern EIA Crest





# All-Sky Imager

- Keo Scientific
- PIXIS 1024 CCD
- 50mm Lens
- 5 Filters
  - 5577
  - 6200
  - 6300
  - 7650
  - 7774
- Ubuntu (Linux) OS







#### Verify a typical lonosphere:

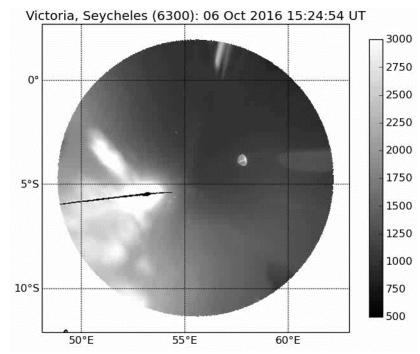
- Pre-reversal enhancement (PRE)
- Ionosphere decay post PRE
- Eastward flow of depletions
- Depletion decay along with lonosphere
- Seasonal trends
  - How does lonosphere visibility change
  - How does depletion frequency/extent change
- Downward trend in Ionosphere visibility and depletions synonymous with solar cycle

Events:

- What's atypical
- Are events correlated by Dst behavior

# Typical Night

- Imager starts shortly after sunset
- Occasional view impacts:
  - Stars
  - Clouds (ITCZ)
  - Moon
  - Rain
  - Planes
  - Lizards
- Ionosphere decays/Brightness decreases through time
- Imager stops shortly before sunrise





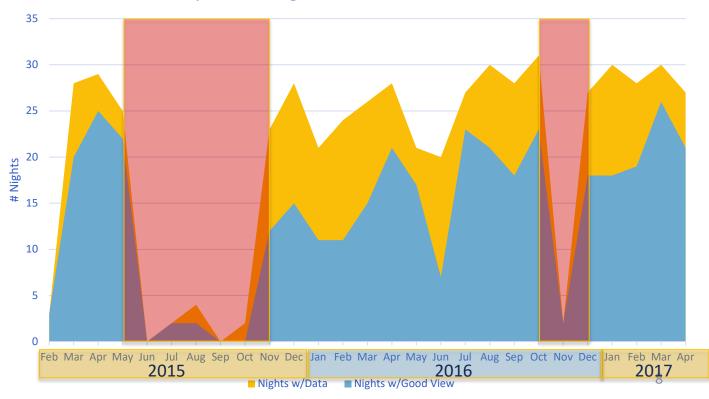
### **Data Collection**

By visual inspection:

- Moon Visibility in PMN or AMN segment.
- Percentage of view obstruction PMN and AMN.
  - 0 = 0%-25% | 1 = 25%-50% | 2 = 50%-75% | 3 = 75%-100%
- Gradient in the ionosphere, in 5577, 6300 and 7774.
- Are there depletions.
- If the depletion was visible in 5577, 6300 or 7774 and if it was PMN or AMN.
- Depletion extent and number of channels.
- EIA starting and ending location.



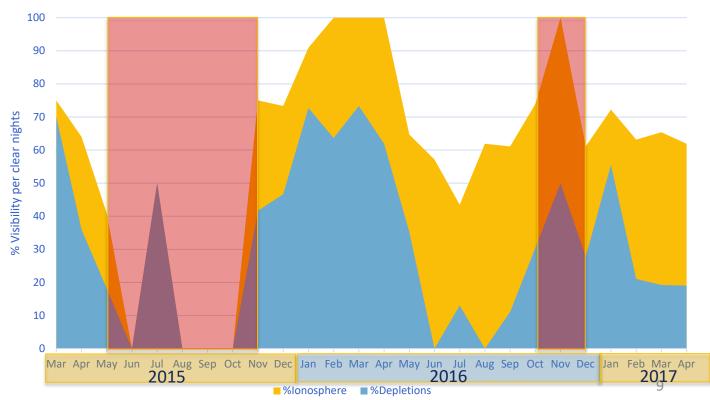
#### Seychelles # Nights w/Data and w/Good View





# Data: Monthly

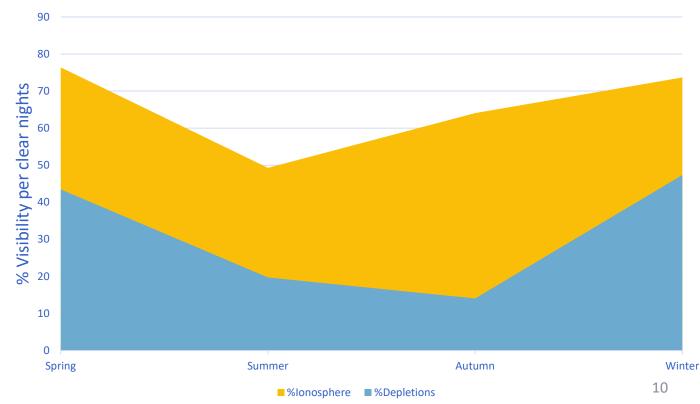
#### Seychelles Ionosphere and Depletion Visibility per Clear Night





### **Data: Seasonal**

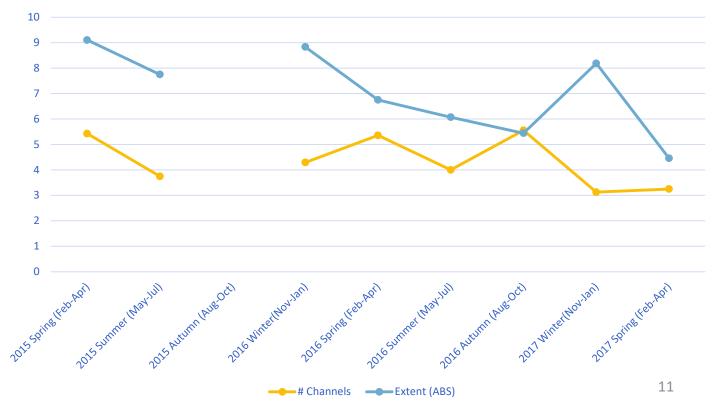
#### Seychelles Ionosphere and Depletion Visibility by Season: All Data





### **Depletion** Data

#### Number of Channels and Latitudinal Extent of Depletions



#### Storm Events

Storm Events:

• March 17<sup>th</sup>, 2015

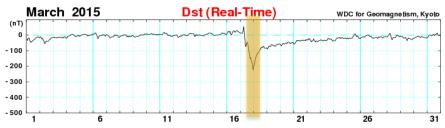


• March 30<sup>th</sup>, 2017

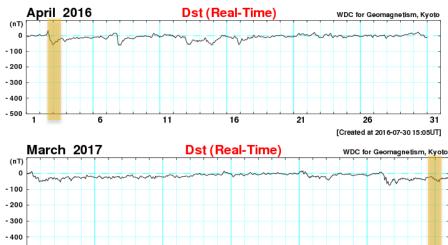
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11



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16

21

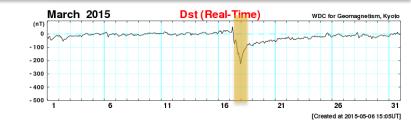
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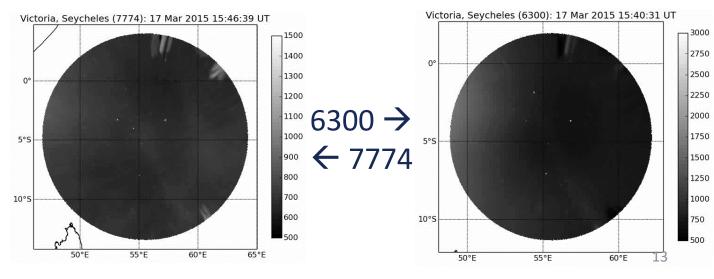
31

26

# March 17th, 2015

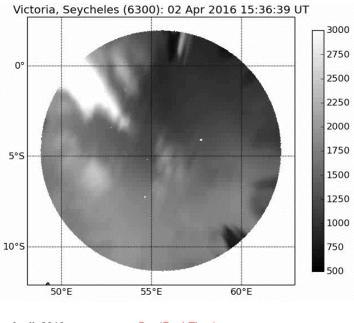
- Strong drop in Dst
- Two large depletion channels
- Westward movement of depletions
- Tilt on southern end of depletions
- Brightest in 6300
- Visible but dimmer in 7774
- One EIA brightening





# April 2nd, 2016

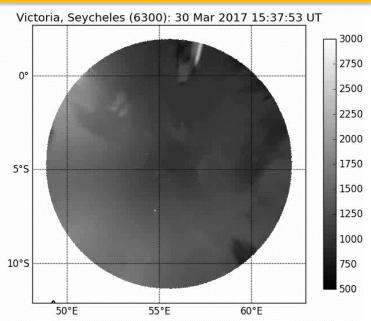
- Weaker drop in Dst
- Eastward and westward movement
- Slight tilt (not as defined as on 3/17/15) in westward transition.
- Brightest in 6300.
- Two peaks in brightness of the EIA





# March 30th, 2017

- No major drop in Dst
- Eastward and westward movement
- Slight tilt in slight westward transition.
- Brightest in 6300.
- Two peaks in brightness of the EIA, not as strong compared to 4/2/16







#### Summary

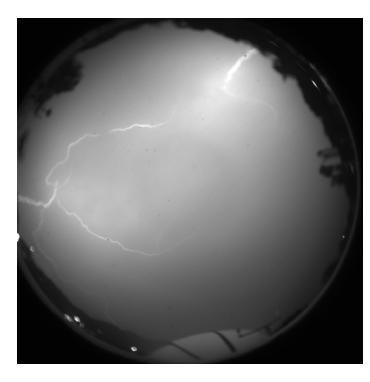
Verifying Ionosphere:

- Pre-reversal enhancement (PRE), Secondary enhancement some nights.
- Ionosphere decay post PRE, some special occasions.
- Night time eastward flow of depletions, but Geomagnetic conditions can reverse direction.
- Depletion degradation along with lonosphere.
- Seasonal similarities:
  - Ionosphere visibility higher in Equinox months.
  - Number of nights with depletions higher in Equinox months.
- Downward trend in Ionosphere visibility and depletions synonymous with solar cycle.
  - Difficult to define due to data outages. Depletion data supports it.
- Events:
- Each event required a sharp or prolonged Dst.
- Can see eastward reversal and westward movement depending on what time you're viewing geomagnetic storm conditions.



### **Acknowledgements**

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