

Imaging of lonospheric Irregularities via Observations of their Impact on the Plasmasphere

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Ionospheric Impact on Interferometers

Observing through plasma

Simultaneously observe cosmic source and ionospheric structure

- Ionospheric delay proportional to total electron content (TEC) along line of sight times v⁻².
- Gradients in TEC lead to additional baseline phase ~δTEC×v⁻¹, so impact much larger at low frequencies (dominates over troposphere blow roughly 1 GHz).

Turning trash into treasure

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- Methods to solve for/remove ionospheric phase provide extremely precise measurement of TEC gradient.
- With VLA low-band system (74 and 330 MHz), δTEC precision as good as 10⁻⁴ TECU (1 TECU = 10¹⁶ e⁻ m⁻²), or gradient precision ~2×10⁻⁴ TECU km⁻¹.

Very Large Array (VLA): Array of 27 parabolic antennas, each 25-m in diameter. In central/western NM





Ionospheric Studies with the VLA

Looking for TIDs, finding flux tubes

Plasmaspheric irregularities

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- Magnetic eastward-directed (MED) waves discovered ٠ with the VLA turned out to be field-aligned irregularities within the plasmasphere.
- Variations in density among flux tubes, combined with co-٠ rotating nature of plasmasphere leads to appearance of wavefronts moving toward magnetic east.





-0.4 -0.2 0.2 0.0 0.4 V_x (km/s) MED waves were prominent during Mar.-Sep. 1990 campaign; appear to move faster than other waves and more prominent at night (Jacobson & Erickson 1992).

0.4

0.2

0.0

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Radial Profile Retrievals

10

10

0.0002

0.0001

0.0000

-0.0001

-0.0002

gradient (TECU km⁻¹)

TEC

Imaging flux tubes

Quasi-tomographic spectral decomposition (QTSD)

- Can exploit co-rotation to reconstruct radial profiles of N_e gradient.
- Apparent phase speed of co-rotating flux tubes is proportional to distance from telescope.
- First step in QTSD process is to decompose time series into modes with Fourier transform (FT) and sliding window.
- Use spatial variation to measure wavenumber vector perpendicular to line of sight; yields distance estimate.
- Isolate those with wavenumber vectors consistent with flux tubes.





01 March 2001 Virgo A observation with VLA at 74 MHz. *Left*: First step in QTSD, spectral decomposition. 1st order TEC spatial derivatives shown, but 2nd order terms analyzed, too. *Above*: Step 2, find those modes with speeds/distances consistent with expected orientation for co-rotating flux tubes.

Radial Profile Retrievals

Imaging flux tubes

Slant range/time images

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- Final step in QTSD is to group modes consistent with co-rotating flux tubes by distance/slant range.
- Summing within slant-range bins and dividing by bin width gives mean N_e gradient profiles as function of time.
- Complementary to method used with Murchison Widefield Array (MWA) in Australia; images flux tubes over large area of sky using position shifts of many sources at ~200 MHz as tracers of TEC gradient in ~2-min. snapshots.



Ionospheric Connection

Mountain waves in space

Flux tubes modified from below

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- Tracing disturbed flux tubes from QTSD map to locations where intersect the ionosphere (300 km altitude) reveals likely origin.
- Located on or to lee side of Rocky Mountains; implies variations in ion pressure from mountain waves (possibly standing waves) in ionosphere responsible for observed density gradient fluctuations in plasmasphere above.
- Note, relatively complex spatial structure implies integrated measurement (e.g., GPS-base TEC) not ideal (recall original time series).

01-MAR-2001



Arrows show surface wind vectors from NOAA North American Regional Reanalysis (NARR).

Synergy with Simulations

altitude

Modeling MED waves

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SAMI2

- SAMI is a model developed by NRL that combines ion production/transport with empirical models for neutral temp./wind and ExB drifts.
- There are 2-D and 3-D versions.
- 2-D version is open source; have modified it to incorporate simple lee wave model to check mountain wave hypothesis for MED wave origin;



Comparison of observed and model plasmaspheric perturbation amplitudes for different field lines and vertical wavelengths.



Time series of electron density horizontal gradient from SAMI2 simulation of one field line per longitude (horizontal wavelength set to 16 km).

SAMI2, λ_v =500 km

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VLA Low-band lonosphere and Transient Experiment (VLITE)

VLITE

Piggybacking on the VLA

- VLITE is dedicated backend on 10 VLA antennas (soon to be more).
- Takes advantage of P-band optics to continually stream 320—384 MHz band to dedicated software correlator (DiFX).
- Each scan output separately; realtime ionospheric pipeline automatically processes these as they arrive.
- Science operations started in Nov. 2014 and still going . . .



P-band feed 4-band feed (no longer used)



Moving Forward

The future

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ABORATORY

Synergy with other platforms

- Beginning to look at overlap with data from other platforms, e.g., coherent backscatter radar (SuperDARN), HF Doppler sounder (LWA1+WWV).
- Could use some help (**subtle hint to grad students/potential postdocs in audience**).

Moving to other platforms

• Exploring development of re-locatable array with low-gain antennas to work in VHF regime where dishes are not needed.





Joppler (Hz)

09-MAR-2015

Schematic of MIT's RAPID system (Lonsdale 2013).

model (and International

Reference lonosphere).



0.0

-1.5

-3.0

-4.5

13.3