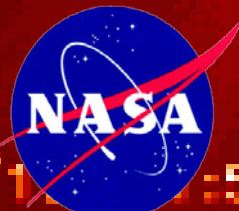


Longitudinal Variability of Nightside Equatorial Electrodynamics?

Endawoke Yizengaw and AMBER Team
Institute for Scientific Research, Boston College, USA



09/11/2011 11:54



Outline

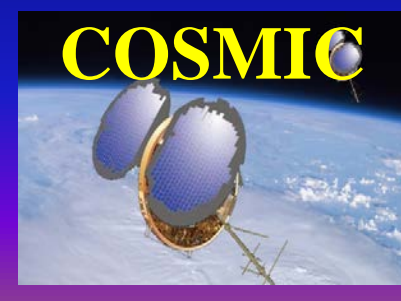
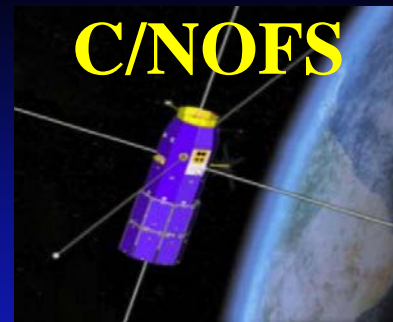
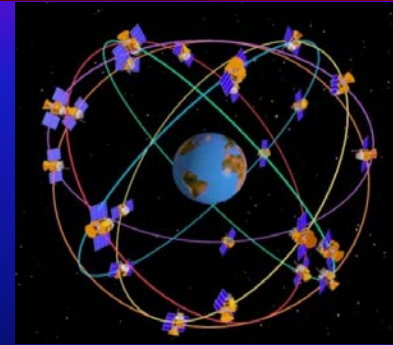
→ Motivation

→ What controls the longitudinal, local time, and seasonal variability of density irregularities?

→ Equatorial electrodynamics?

→ One-to-one correlation between irregularities and vertical drift

→ Future Direction: **Can we use magnetometer to estimate the nightside drifts?**



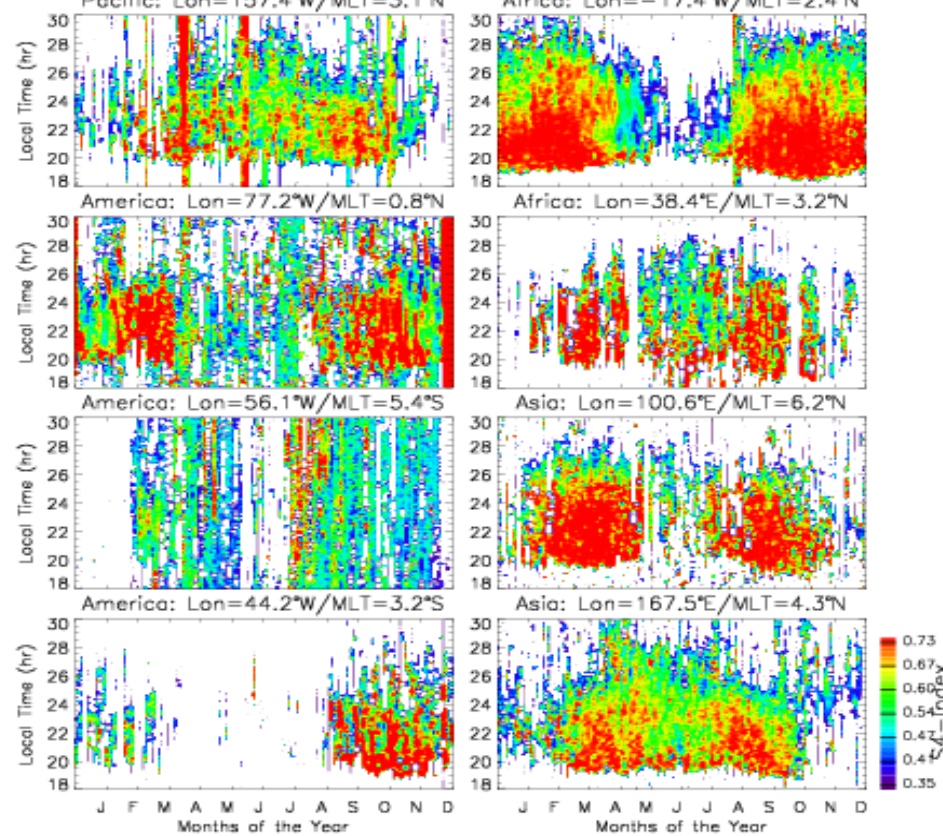
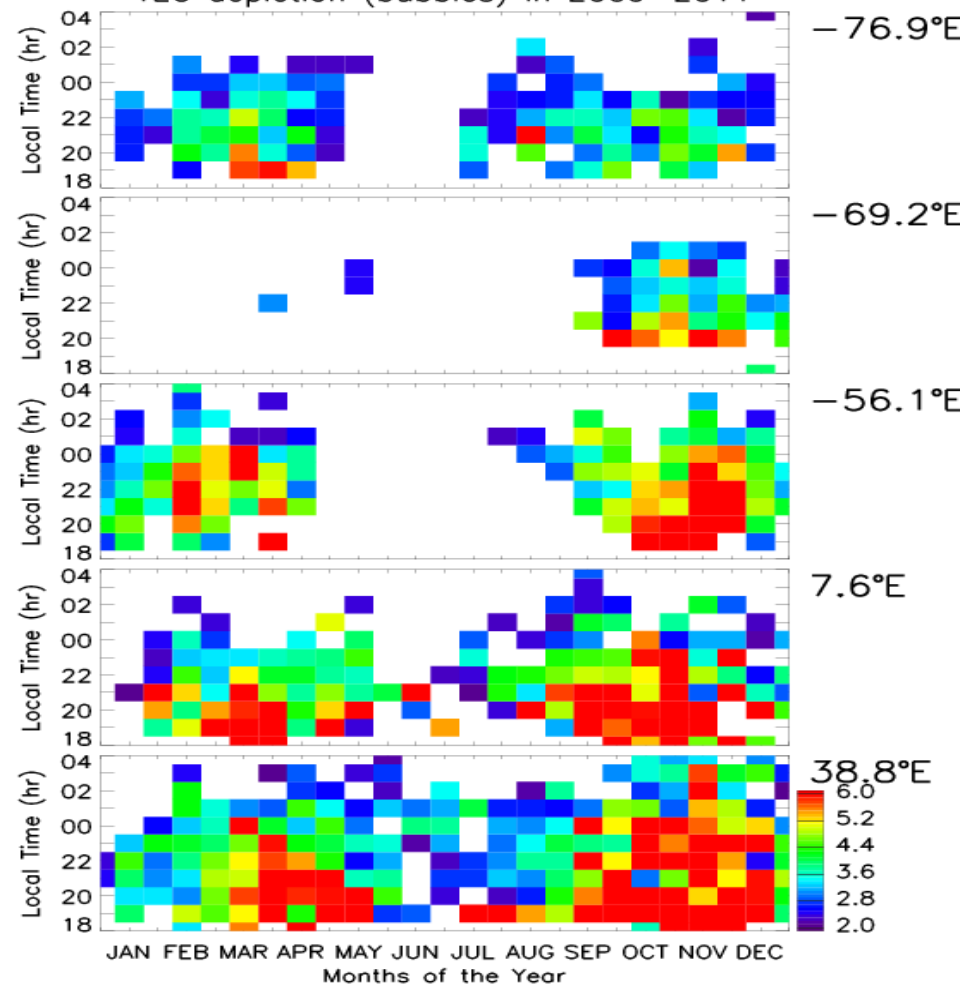
What controls the longitudinal, Local Time and Seasonal variability of ionospheric irregularities?

Bubbles extracted from TEC

Scintillations from UHF receivers

TEC depletion (bubbles) in 2009–2011

Statistical post-midnight bubbles Longitudinal Dependence (2010–2016)

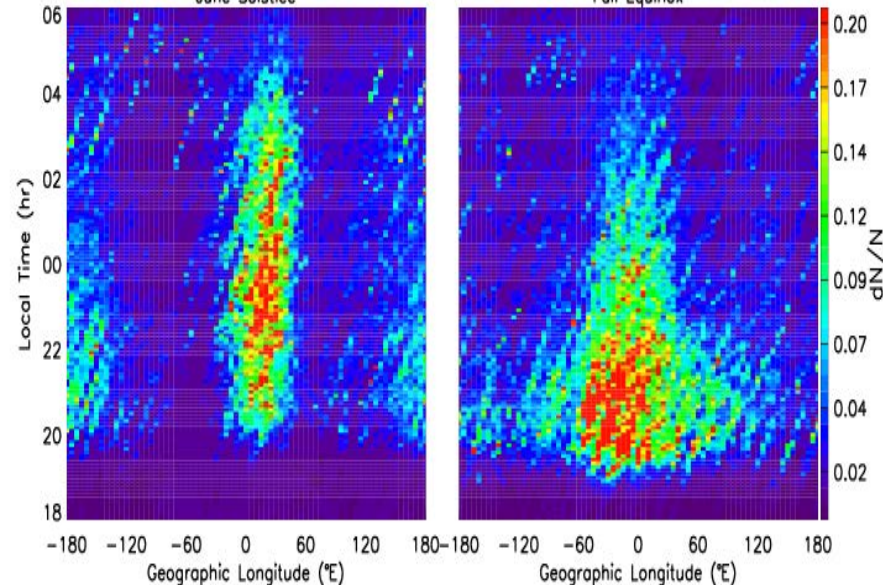
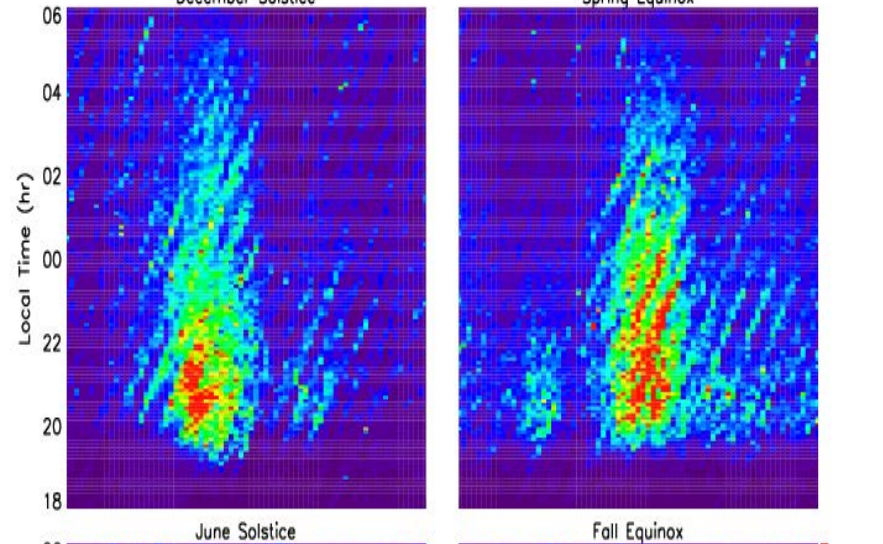


Yizengaw et al., 2017

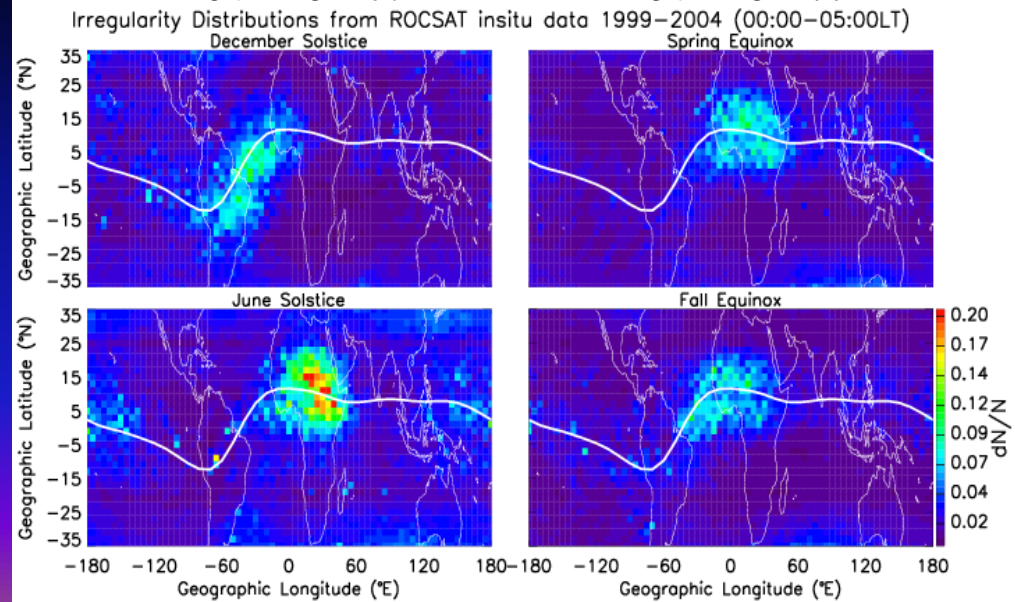
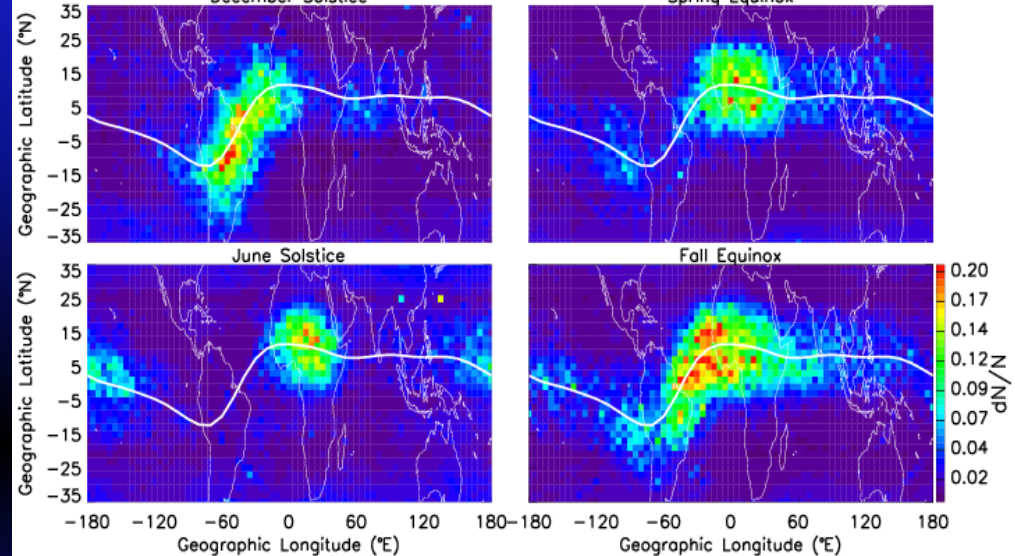
Yizengaw et al., AG, 2014

Longitudinal and seasonal variability of ionospheric irregularities from ROCSAT

Irregularity Distributions from ROCSAT in-situ data 1999–2004
December Solstice Spring Equinox

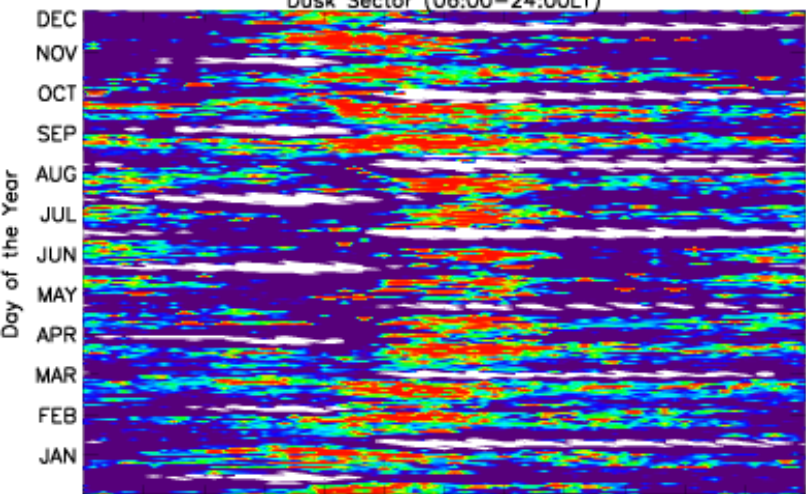


Irregularity Distributions from ROCSAT insitu data 1999–2004 (18:00–24:00LT)
December Solstice Spring Equinox

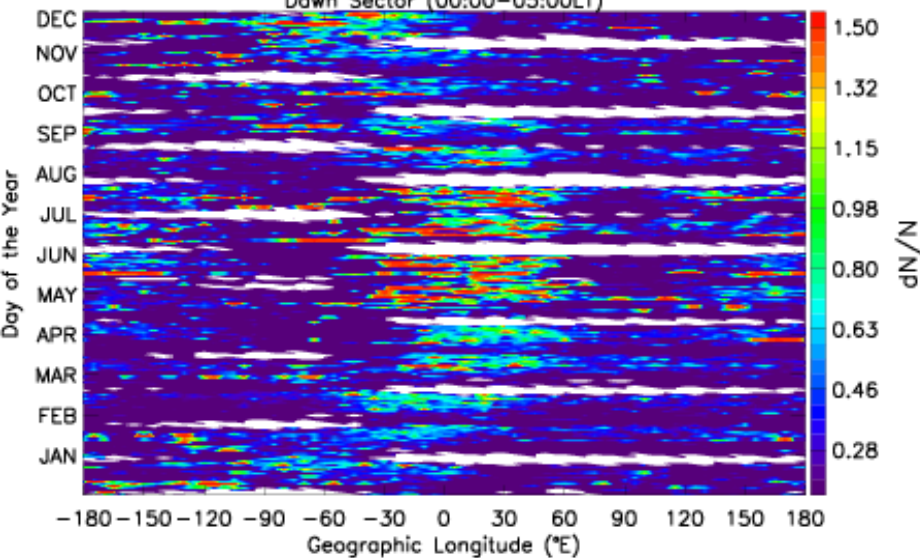


Multi-instrument observations of Longitudinal and seasonal variability of irregularities

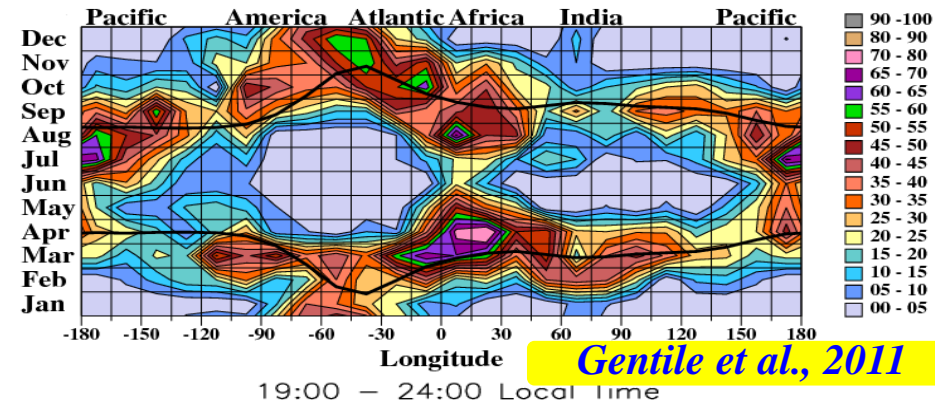
Irregularity Distributions from ROCSAT in-situ data 1999–2004
Dusk Sector (06:00–24:00LT)



Dawn Sector (00:00–05:00LT)

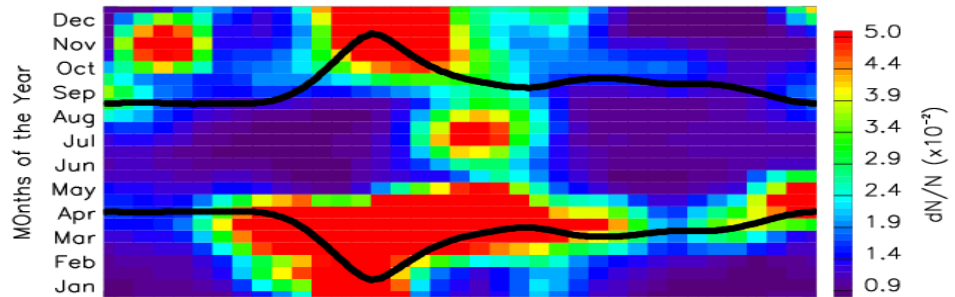


DMSP F15 Evening Sector EPBs 2000

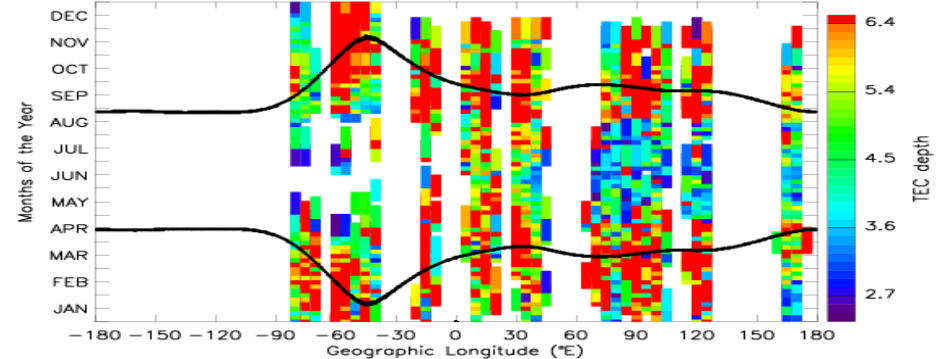


19:00 – 24:00 Local Time

Gentile et al., 2011

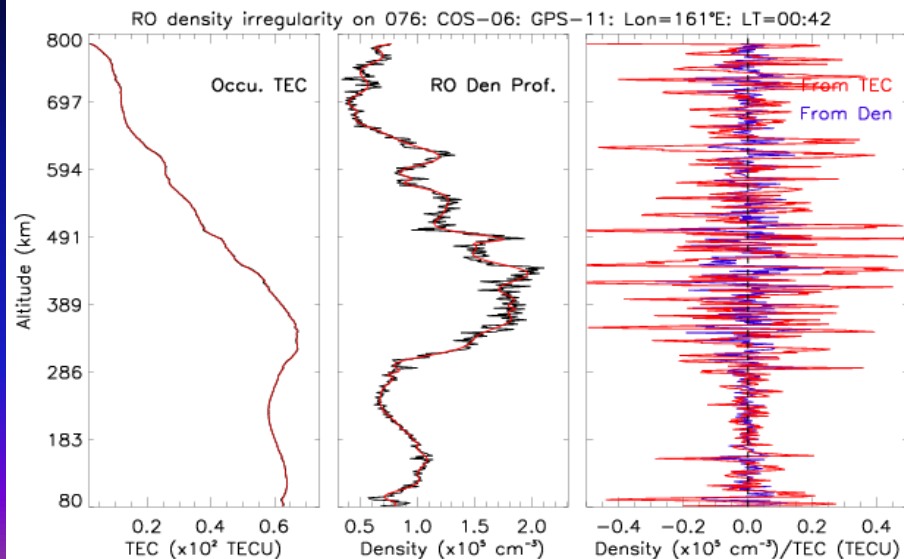
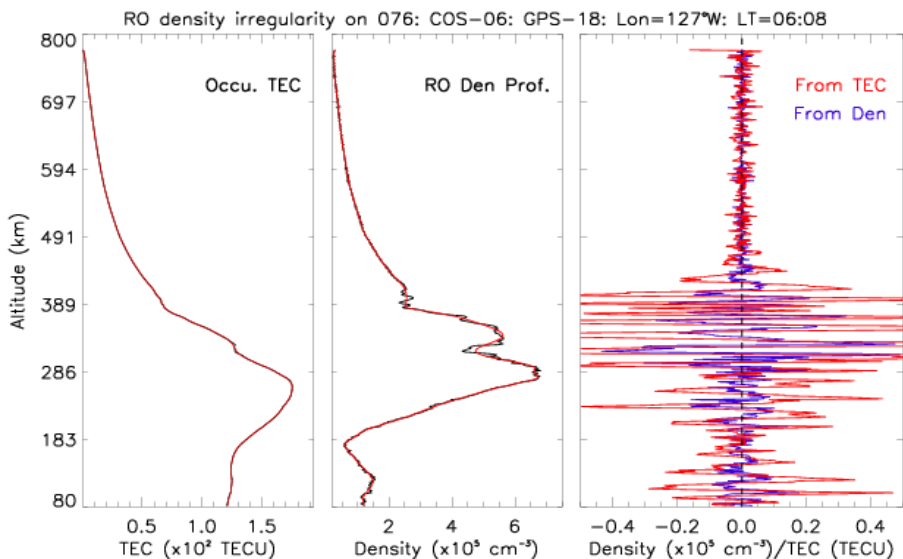
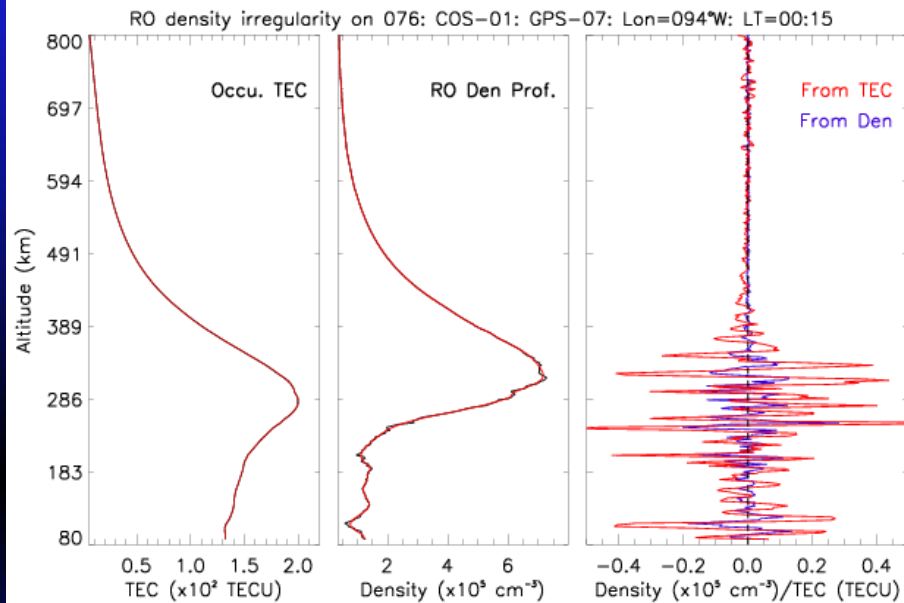
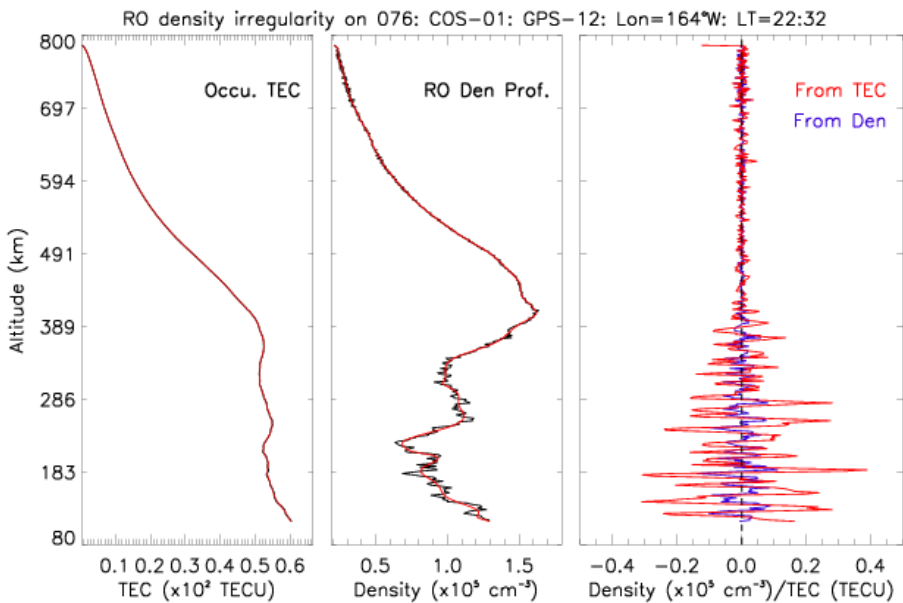


Longitudinal TEC depletion in 2012 (19:00–24:00 LT)



Yizengaw et al., 2014

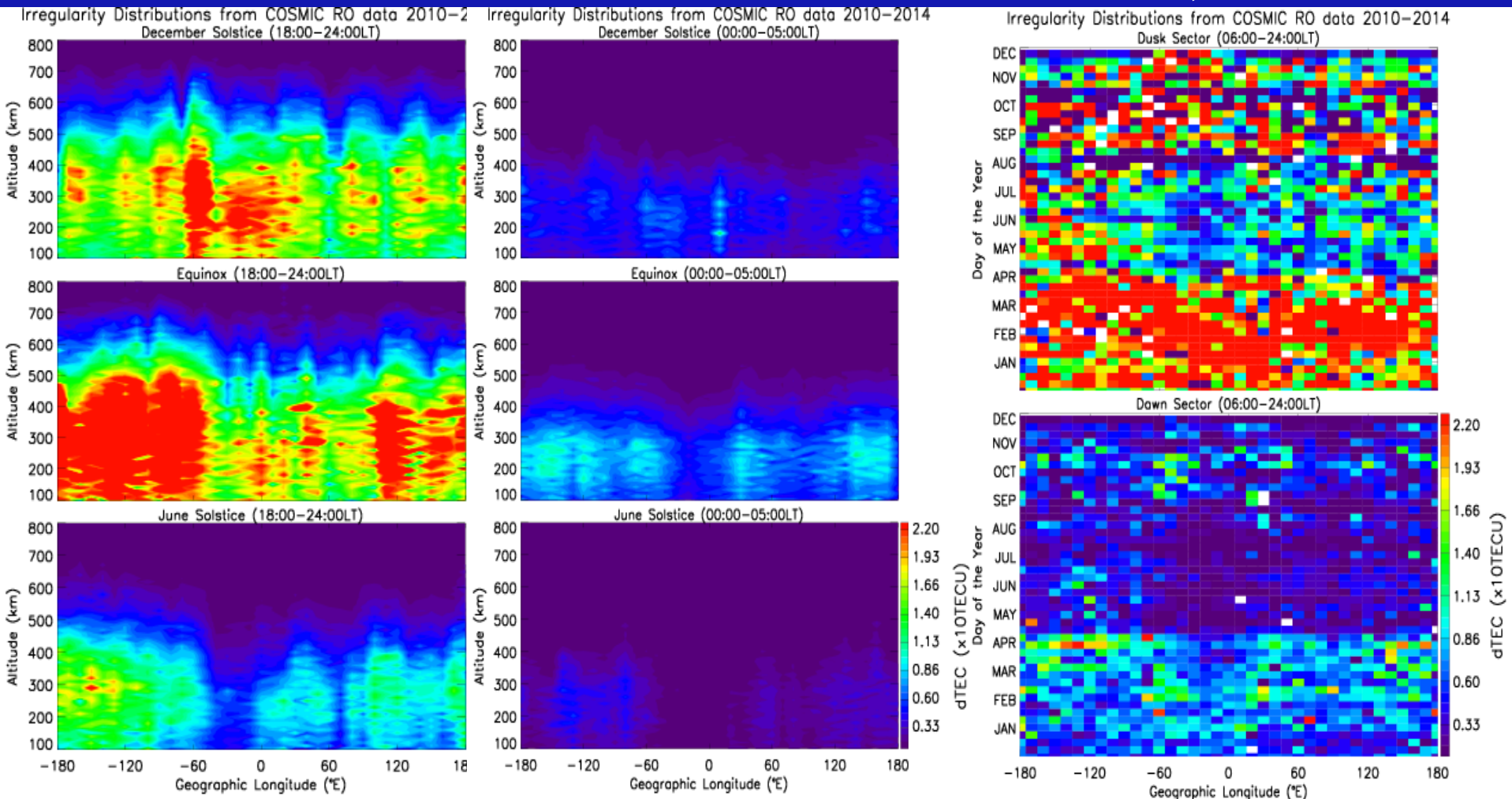
Altitudinal, Longitudinal and seasonal structures of irregularities: Utilizing COSMIC data



Altitudinal, Longitudinal and seasonal structures of irregularities: Utilizing COSMIC data

$Lat = -15^{\circ}$ to $15^{\circ}N$

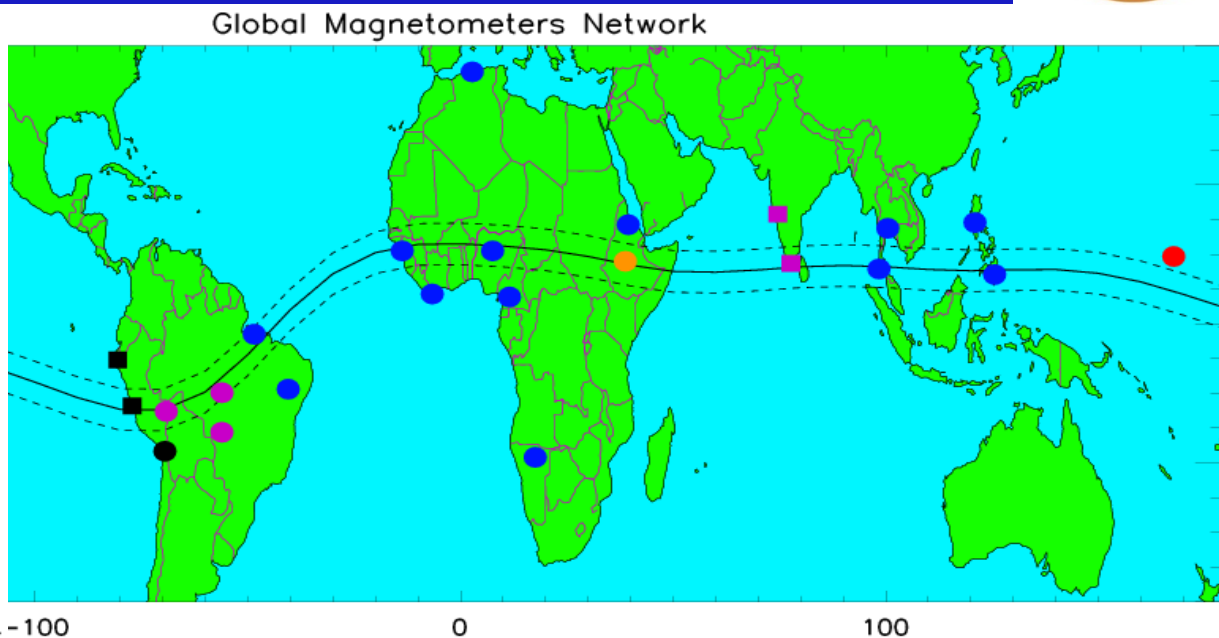
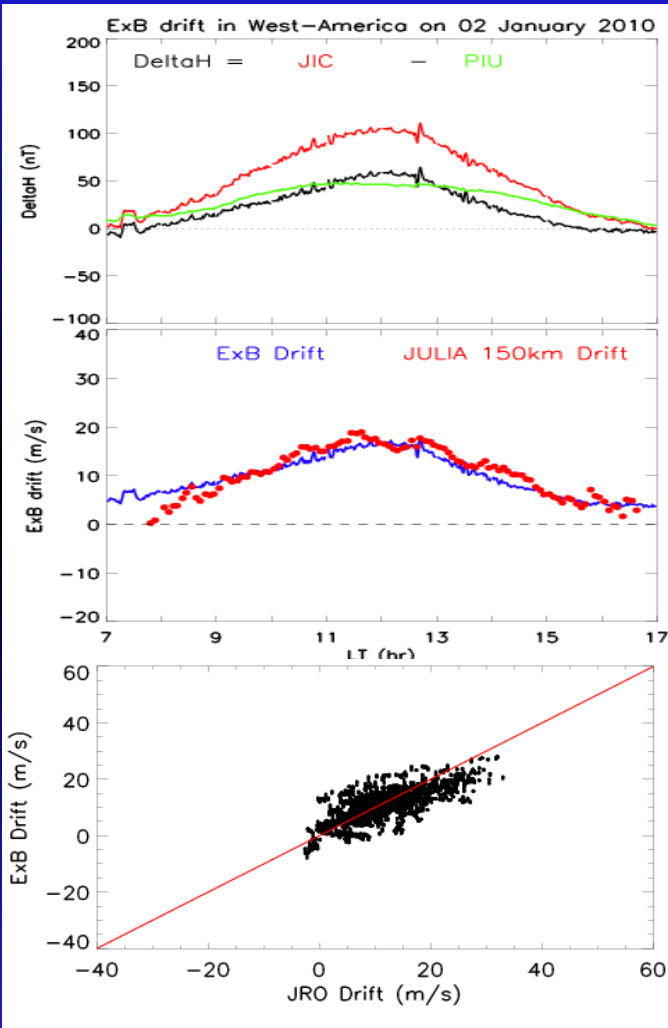
$Lat = -15^{\circ}$ to $15^{\circ}N$; Alt = 200-300km



➔ What controls these variability of irregularities distributions?

**What causes such strong
Longitudinal, Altitudinal, and
Seasonal Variability of the density
irregularities? Is it electrodynamics
or something else?**

Equatorial Electrodynamics: Dayside



➔ Magnetometer at off the equator

$$\mathbf{B}_{\text{Obs}} = \mathbf{B}_{\text{main}} + \mathbf{B}_{\text{SQ}} + \mathbf{B}_{\text{FAC}} + \mathbf{B}_{\text{RC}} + \mathbf{B}_{\text{MP}}$$

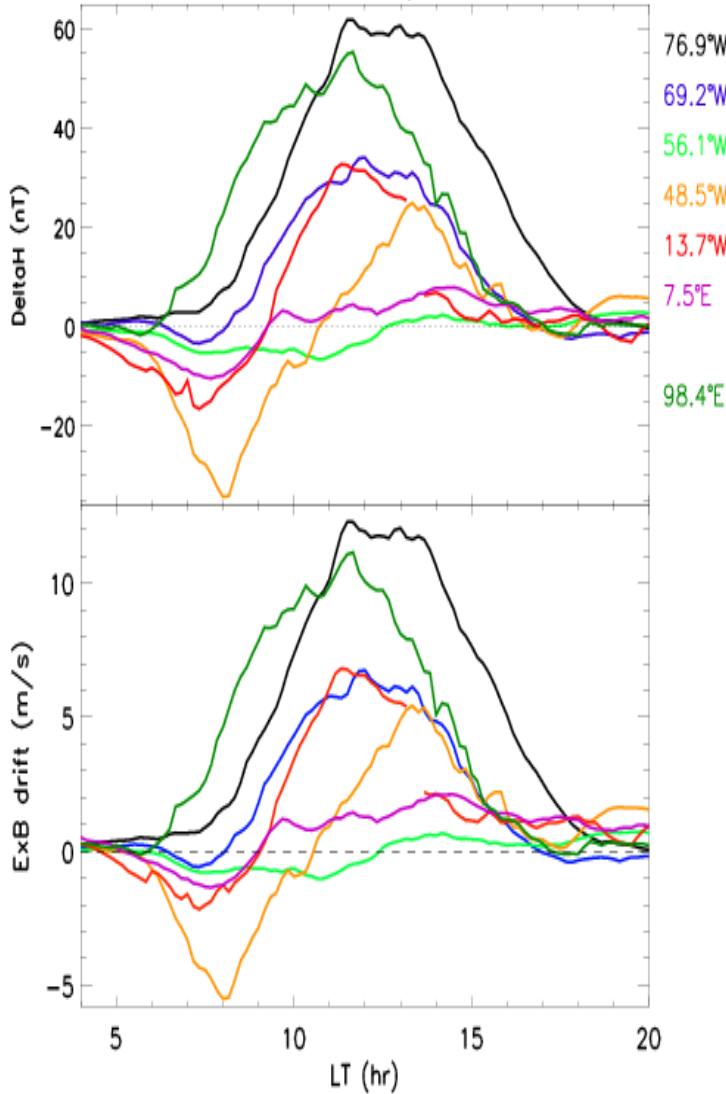
➔ Magnetometer at the equator

$$\mathbf{B}_{\text{Obs}} = \mathbf{B}_{\text{main}} + \mathbf{B}_{\text{SQ}} + \mathbf{B}_{\text{FAC}} + \mathbf{B}_{\text{RC}} + \mathbf{B}_{\text{EJ}} + \mathbf{B}_{\text{MP}}$$

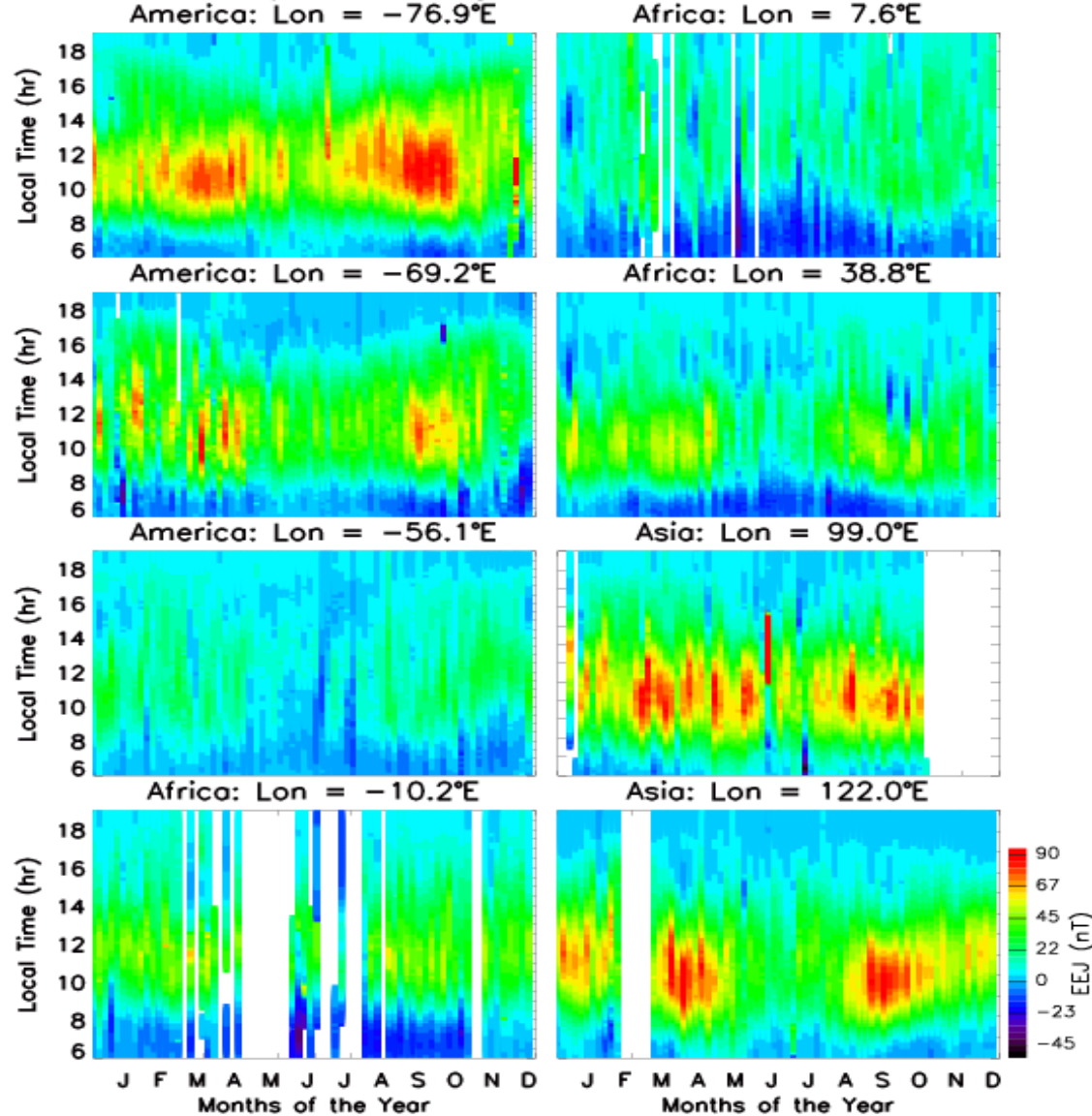
➔ **Problem:** Does EEJ current or Equatorial Electrodynamics in general show longitudinal and seasonal dependence?

Longitudinal variability of dayside drift

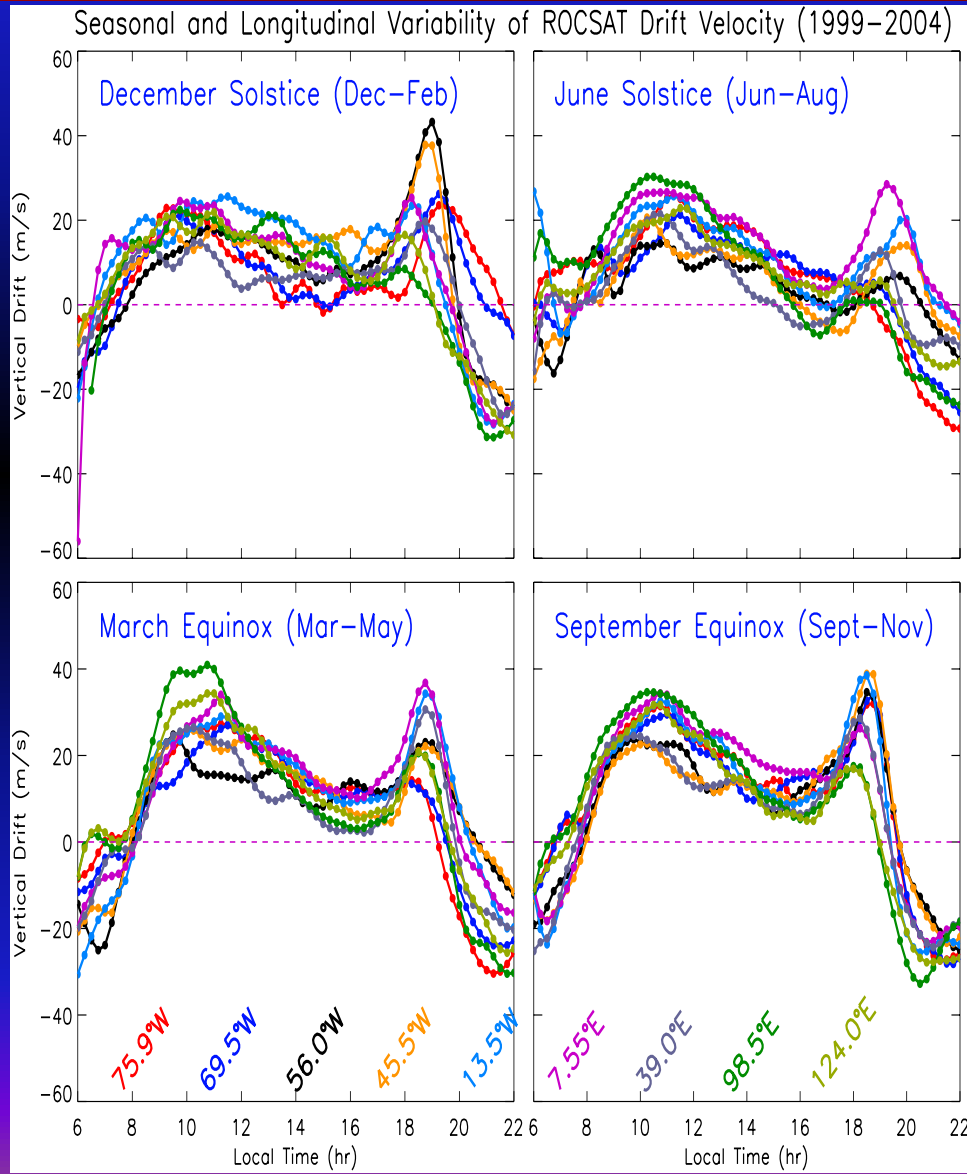
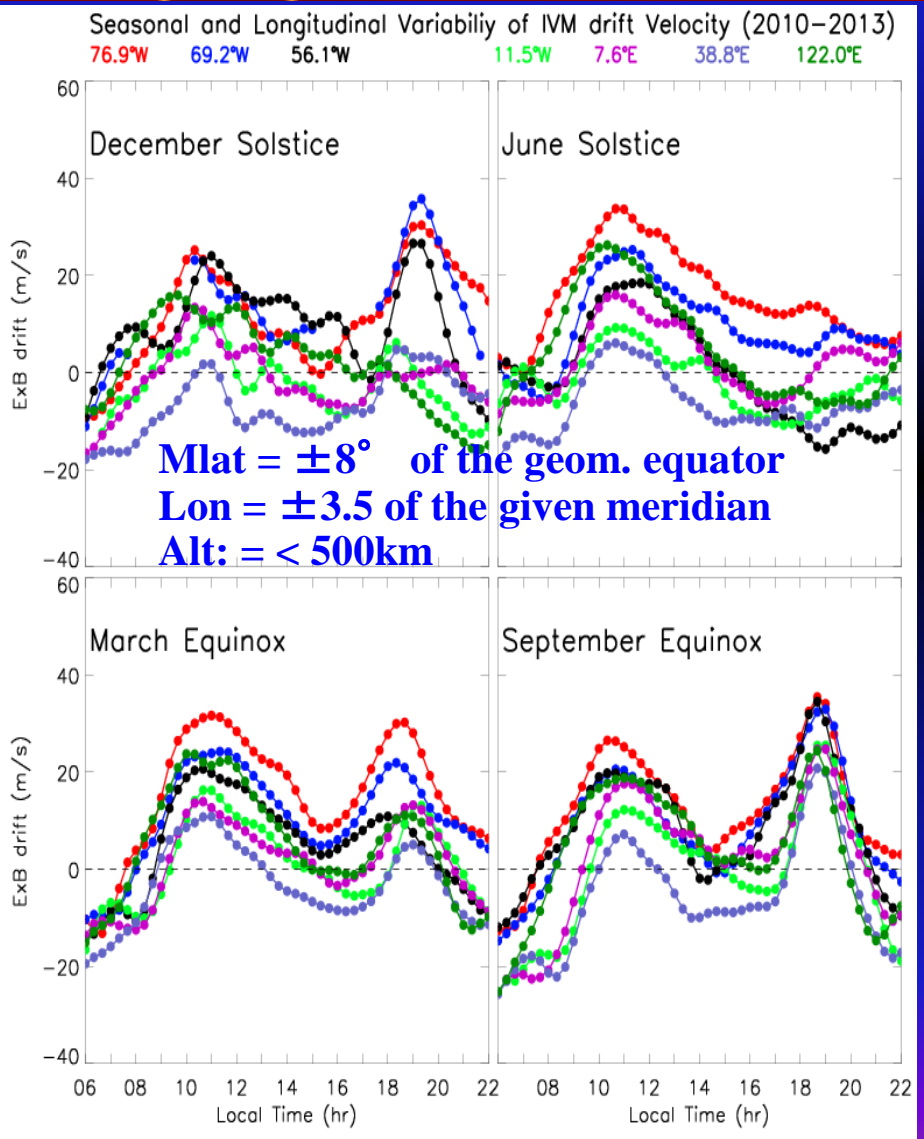
ExB drift in July 2015



Statistical dayside EEJ Longitudinal Dependence (2010–2015)



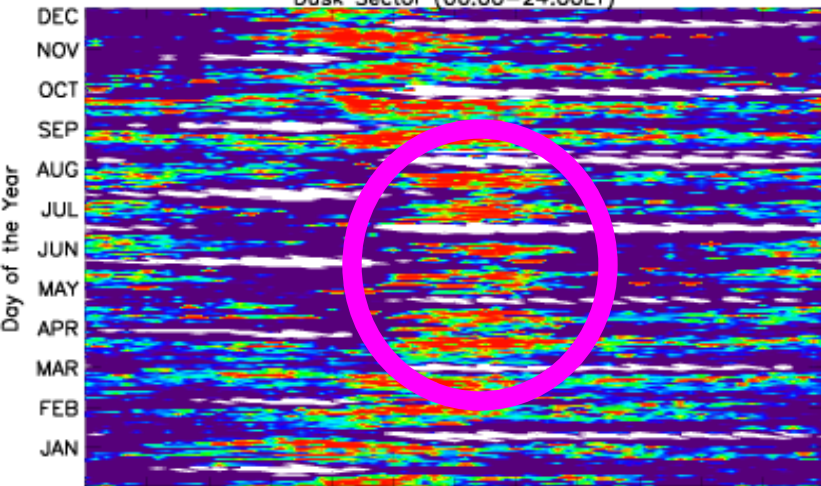
Longitudinal variability of day and night side drifts



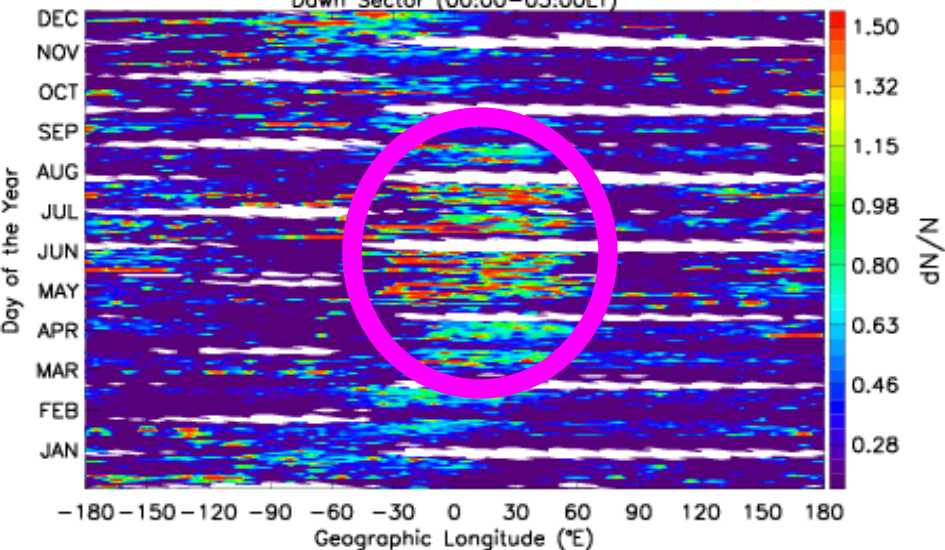
One-to-one correlation between Longitudinal and Seasonal variability of drifts and irregularities

Irregularities at ROCSAT alt.

Irregularity Distributions from ROCSAT in-situ data 1999–2004
Dusk Sector (06:00–24:00LT)

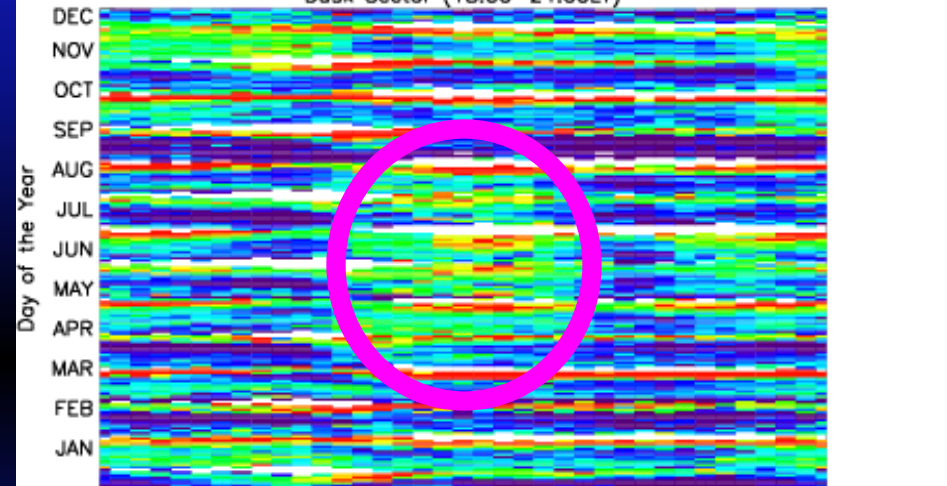


Down Sector (00:00–05:00LT)

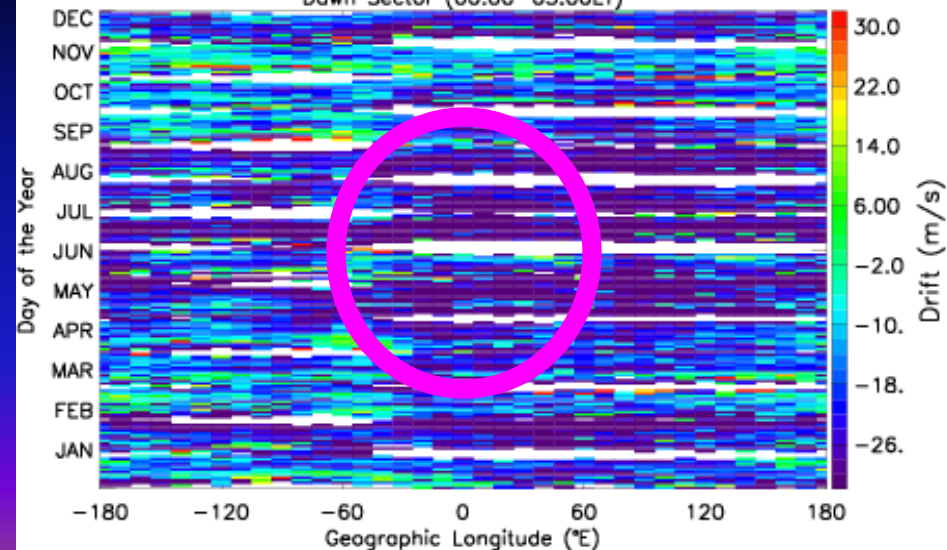


Drift at ROCSAT alt

Equatorial Vertical Drift from ROCSAT insitu data 1999–2004
Dusk Sector (18:00–24:00LT)



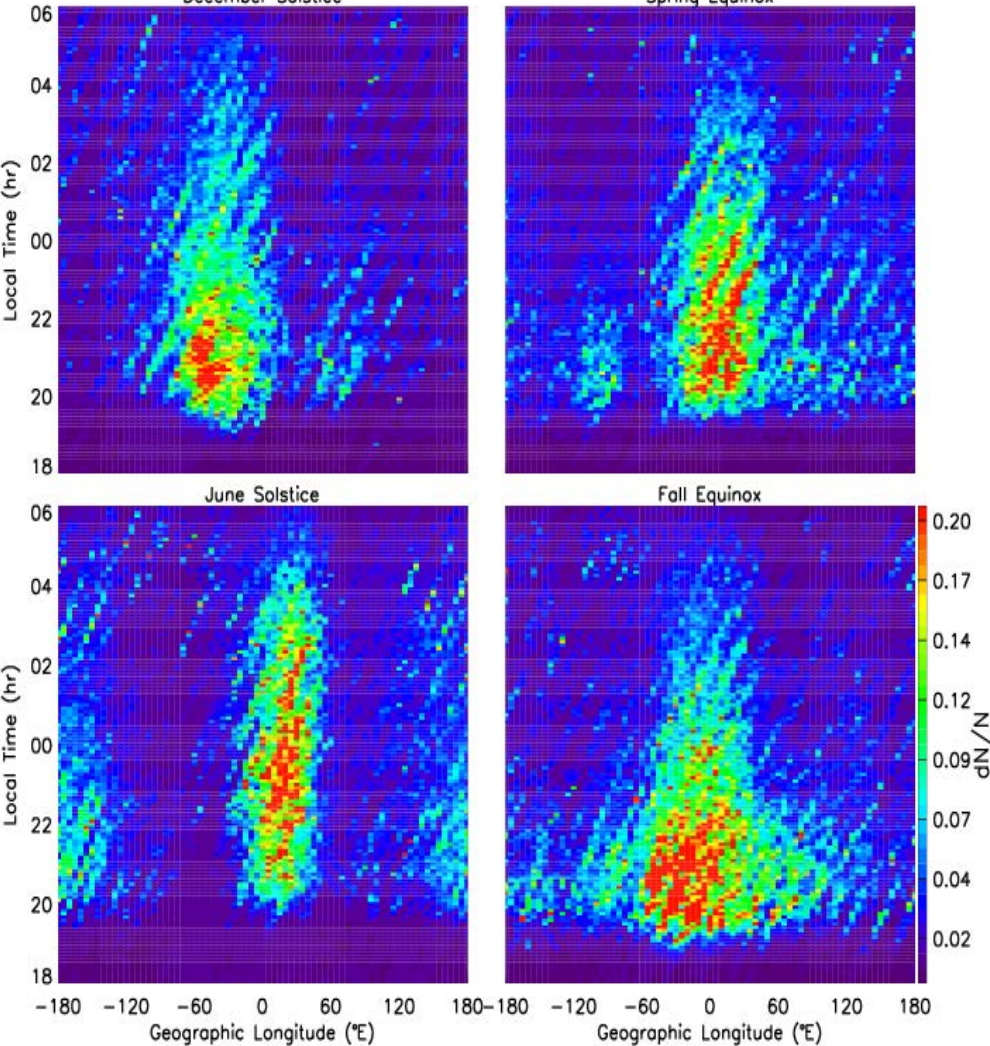
Down Sector (00:00–05:00LT)



One-to-one correlation between Longitudinal, Local time and Seasonal variability of drifts and irregularities

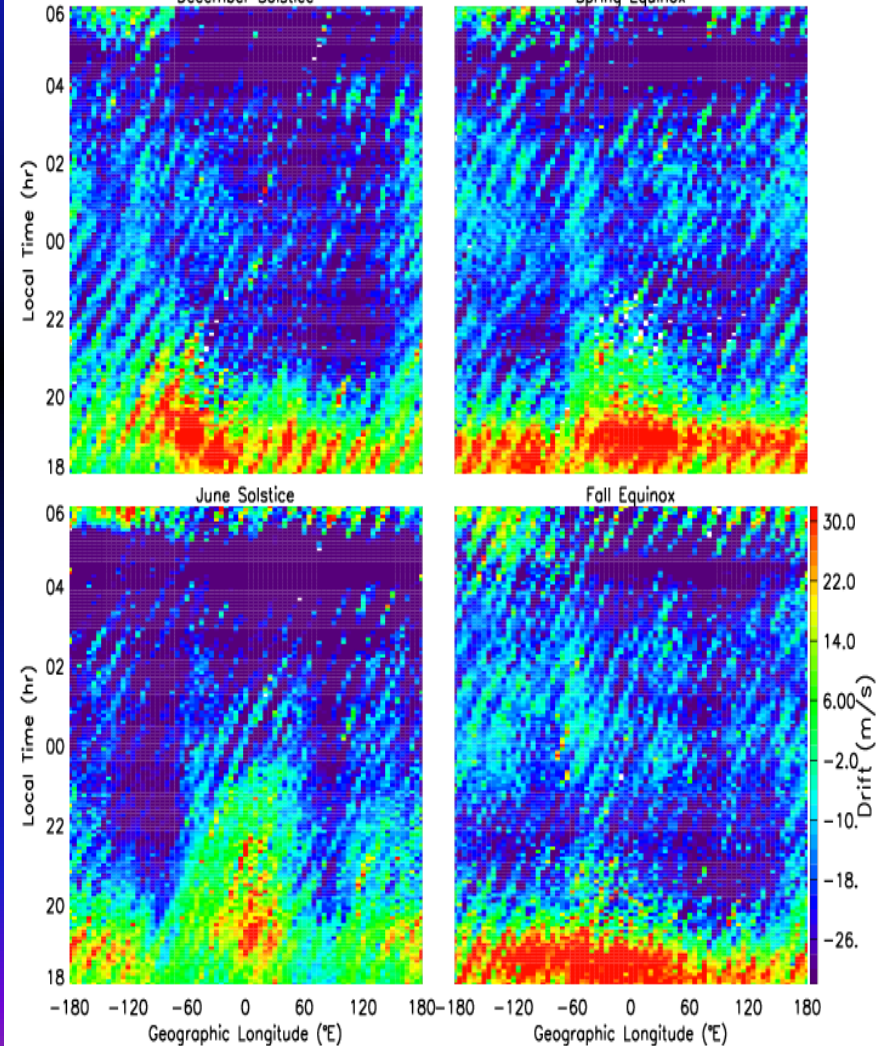
Irregularities at ROCSAT alt.

Irregularity Distributions from ROCSAT in-situ data 1999–2004
December Solstice Spring Equinox

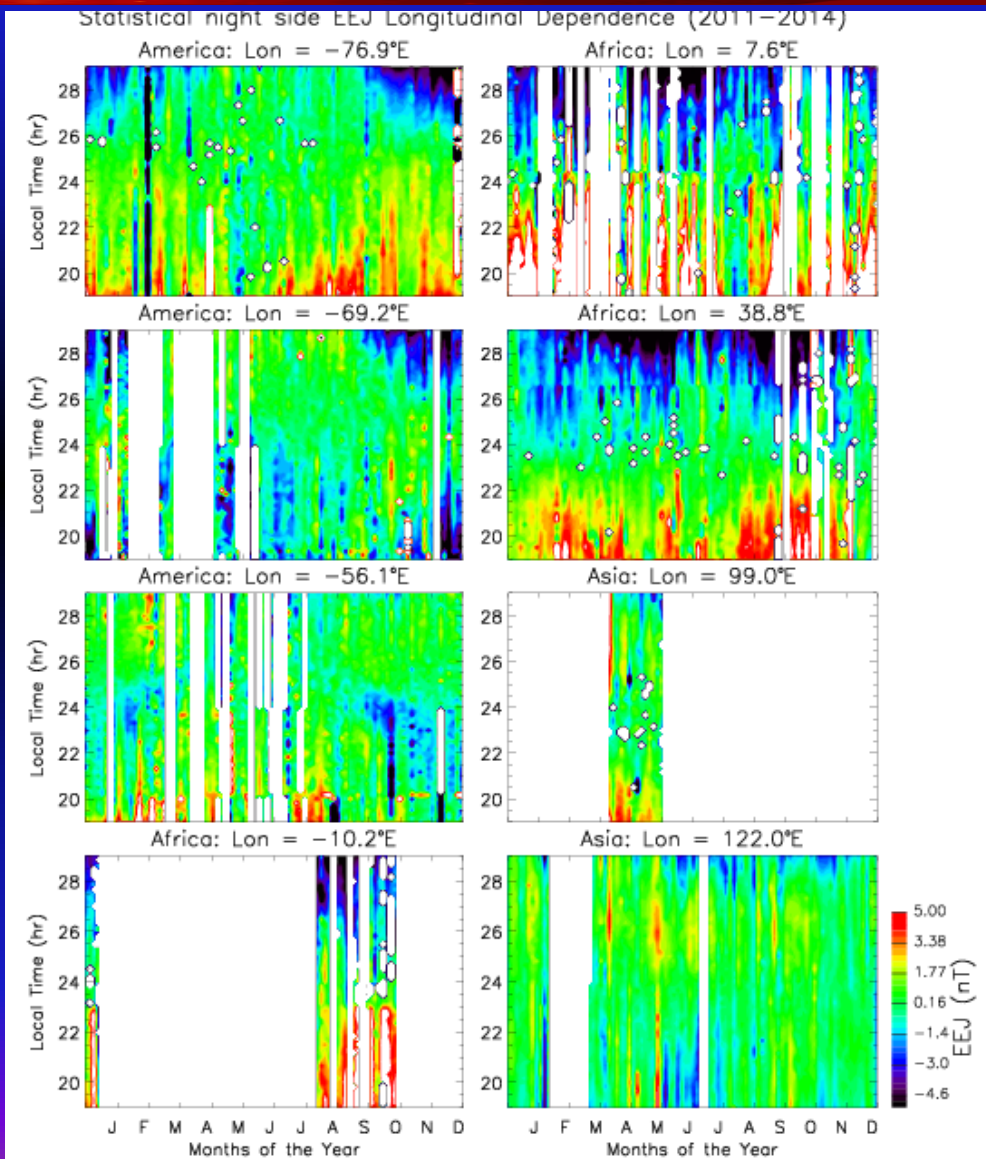
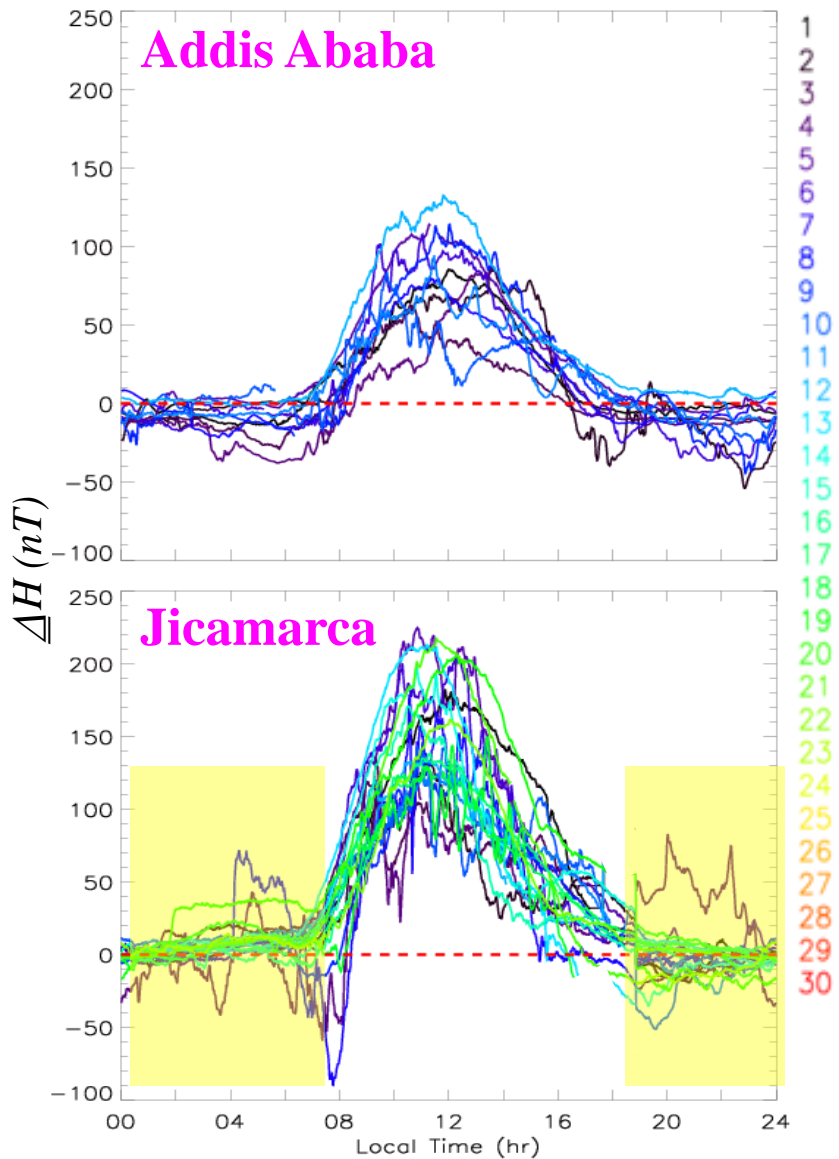


Drift at ROCSAT alt

Equatorial Vertical Drift from ROCSAT insitu data 1999–2004
December Solstice Spring Equinox

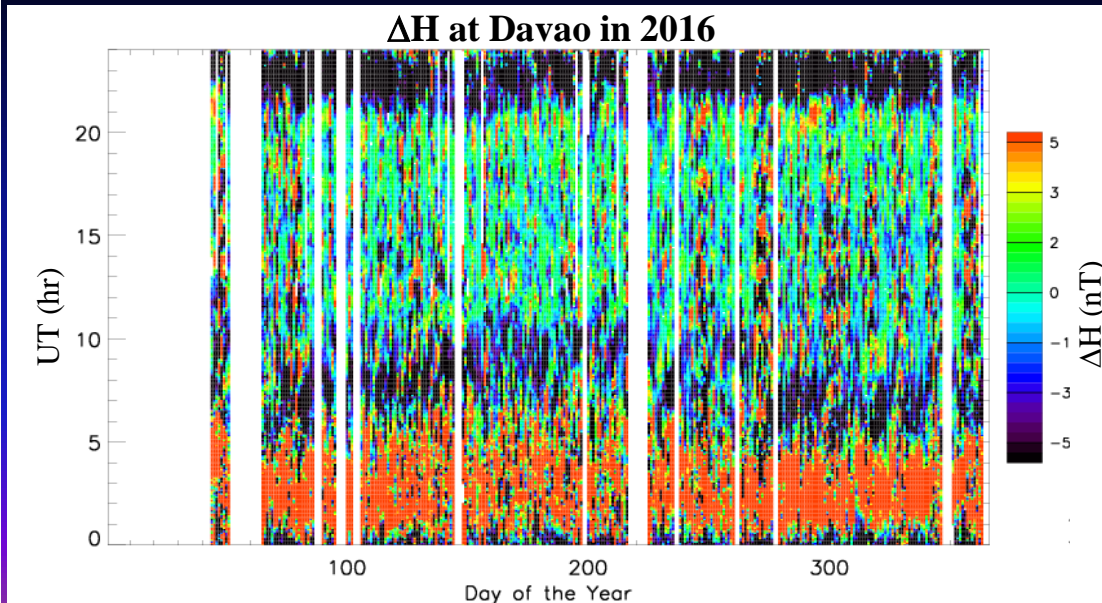
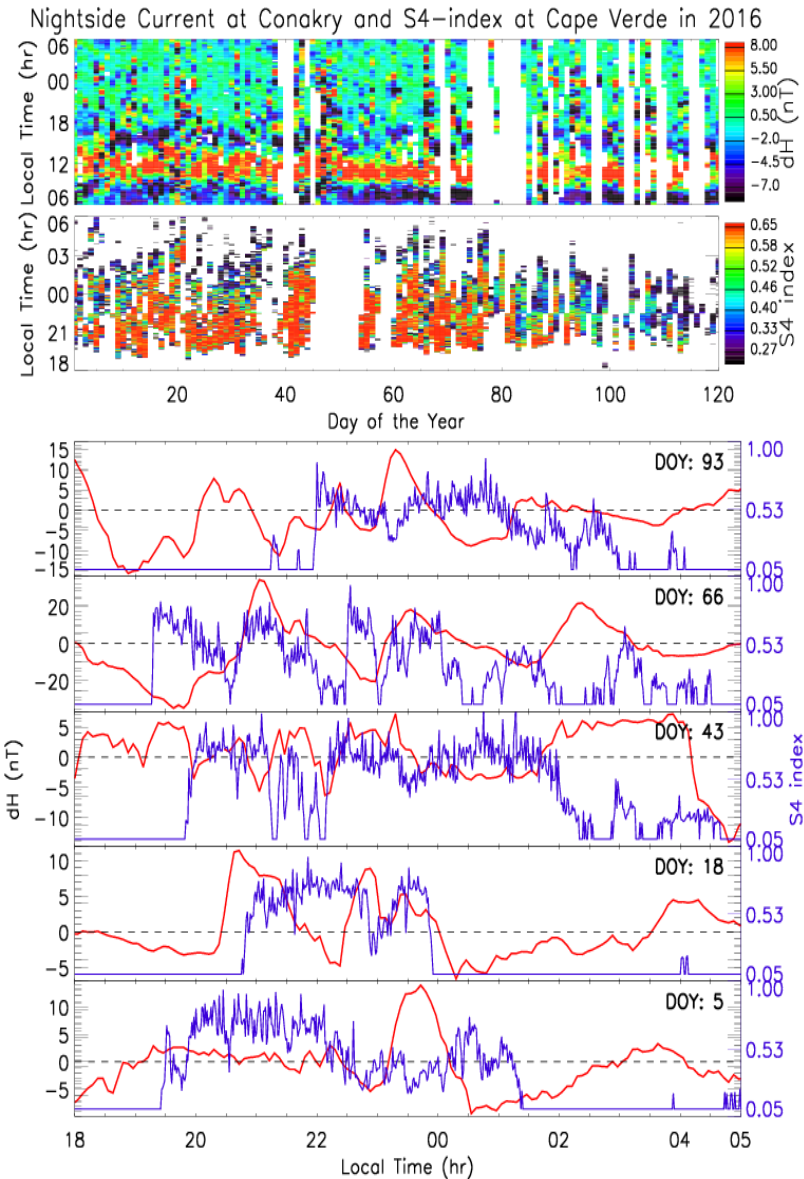


Does Mag data useful to detect the day-to-day variability of nightside electrodynamics?



Does Mag data useful for day-to-day variability of nightside electrodynamics?

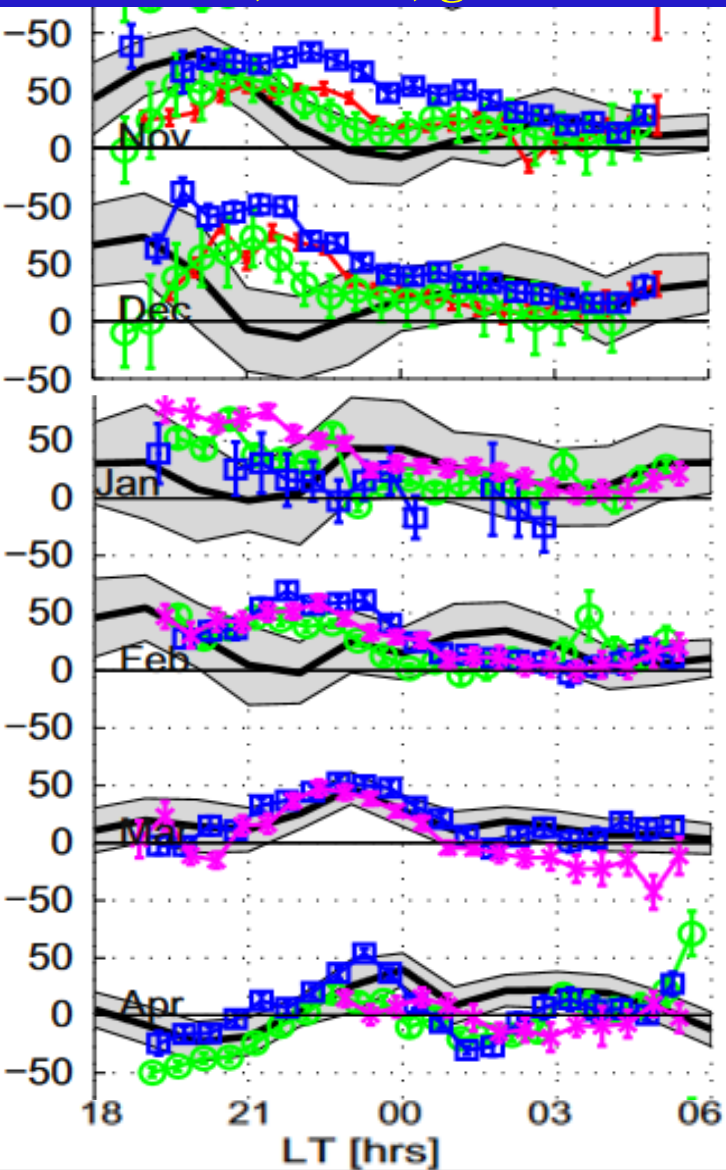
- To reduce the ring current contribution and to removes the gradual drift of the background field, a 4 hours running mean is subtracted from the H -values.
- The residual (ΔH) variations indicates the east-west current or electric field variations during dayside and nightside.



Meridional wind Longitudinal Variability

Brazil 7°S, 39°W; geom: 5.8°S

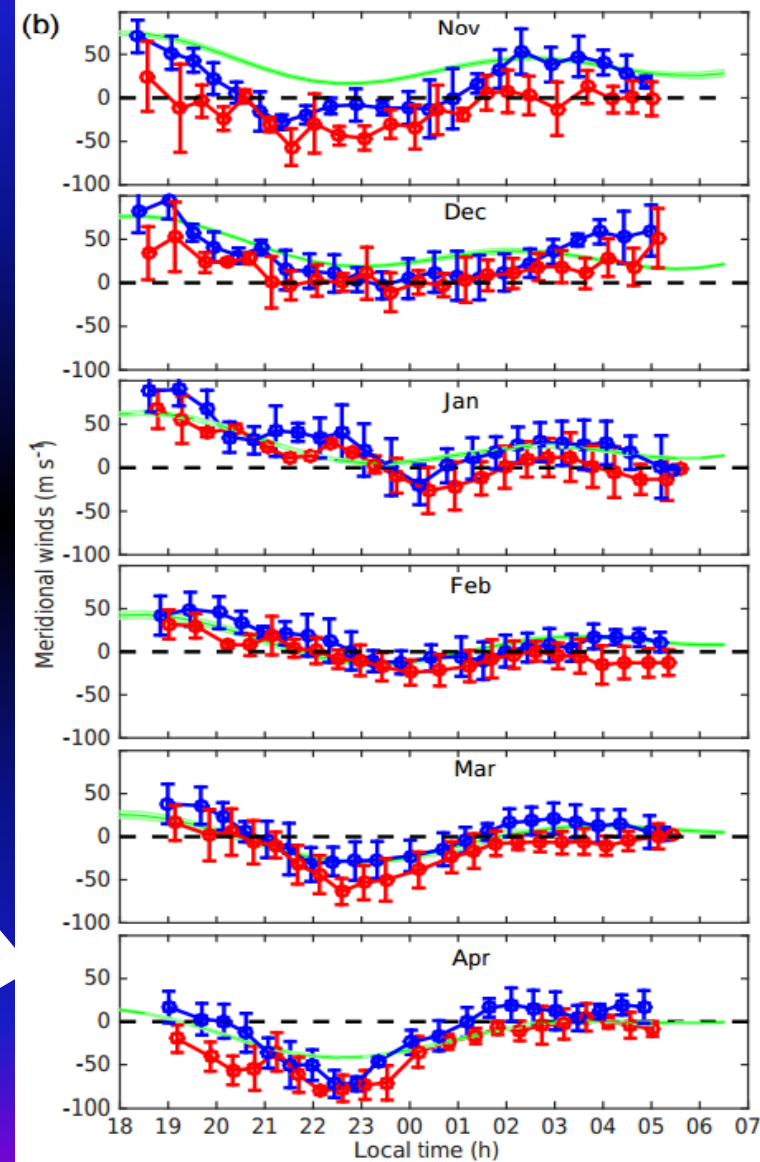
East-Africa 11.6°N, 37.4°E; geom: 3.7°N



2009 2010
2011 2012
WAM model

2015 - 2016
South direction
North direction
HWM14 model

Meriwether et al., JASTP, 2013

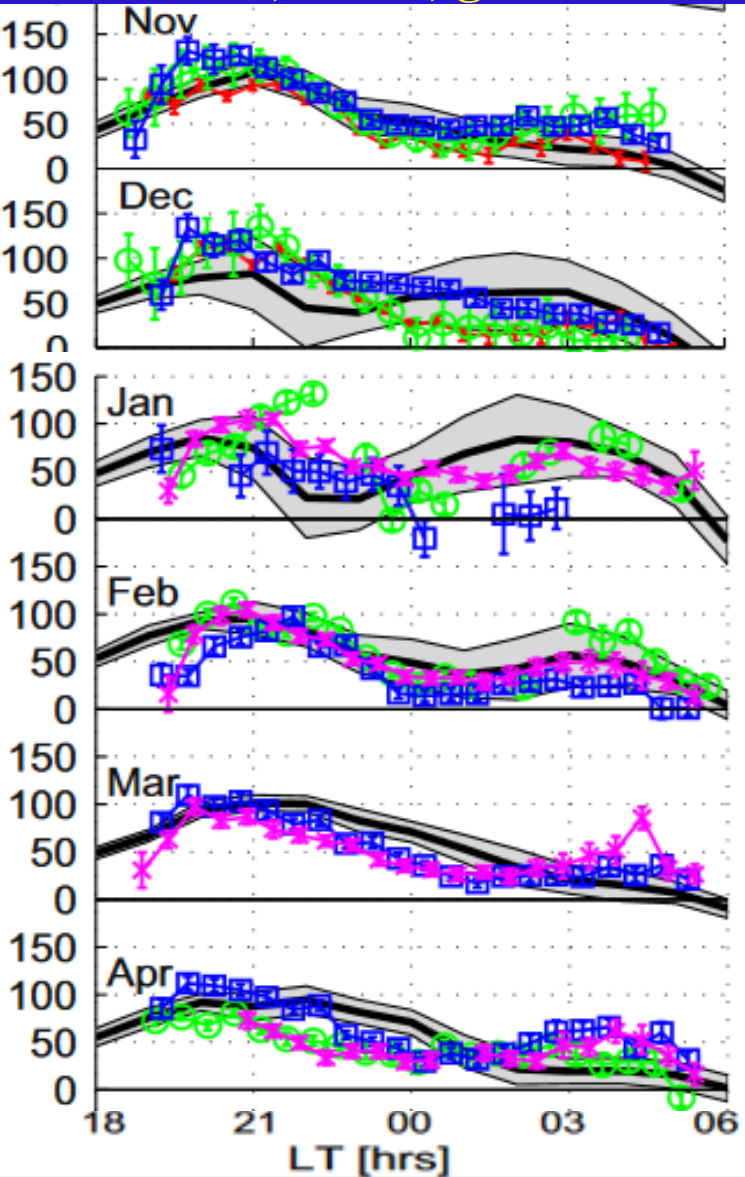


Tesema et al., AG, 2017

Zonal wind Longitudinal Variability

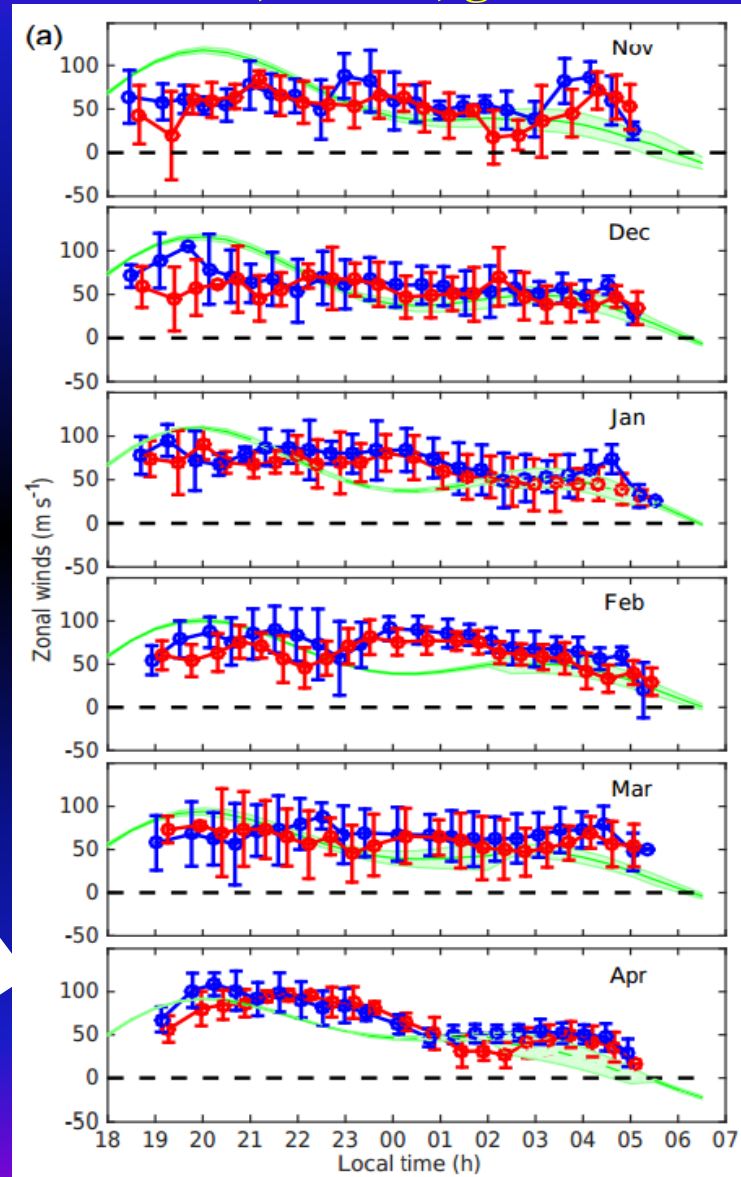
Brazil 7°S, 39°W; geom: 5.8°S

East-Africa 11.6°N, 37.4°E; geom: 3.7°N



2009 2010
2011 2012
WAM model

2015 - 2016
West direction
East direction
HWM14 model

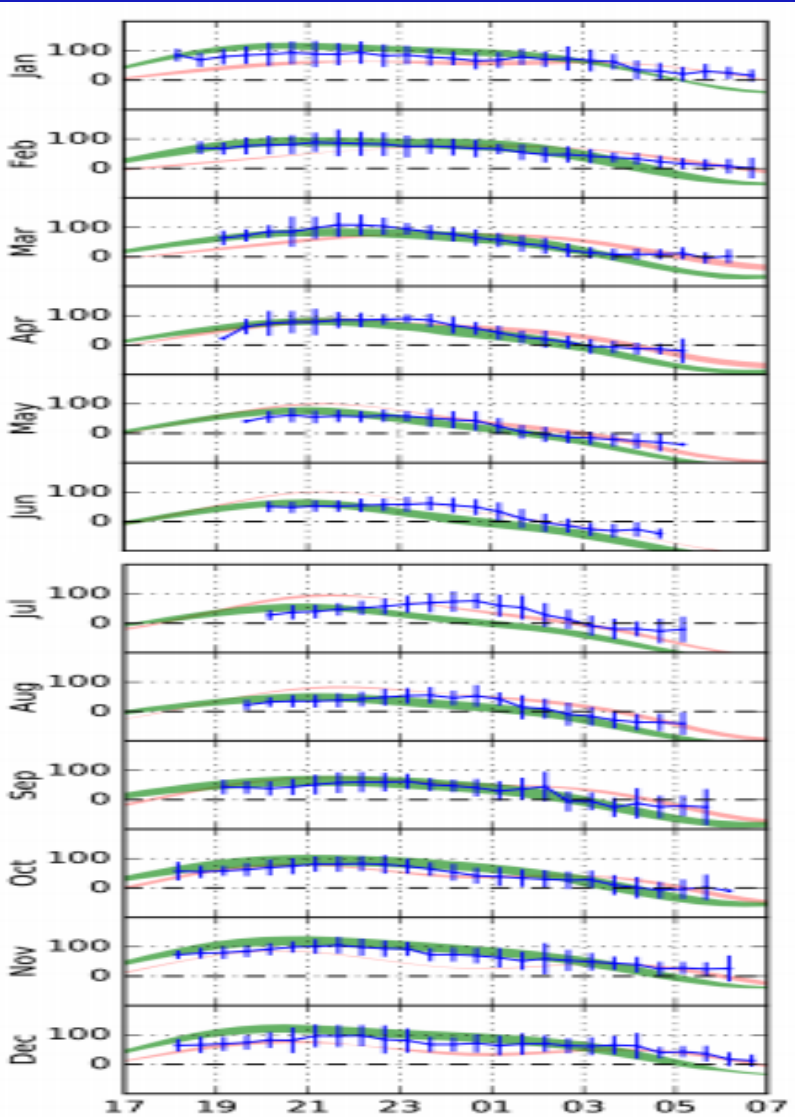


Zonal and Meridional winds in Morocco

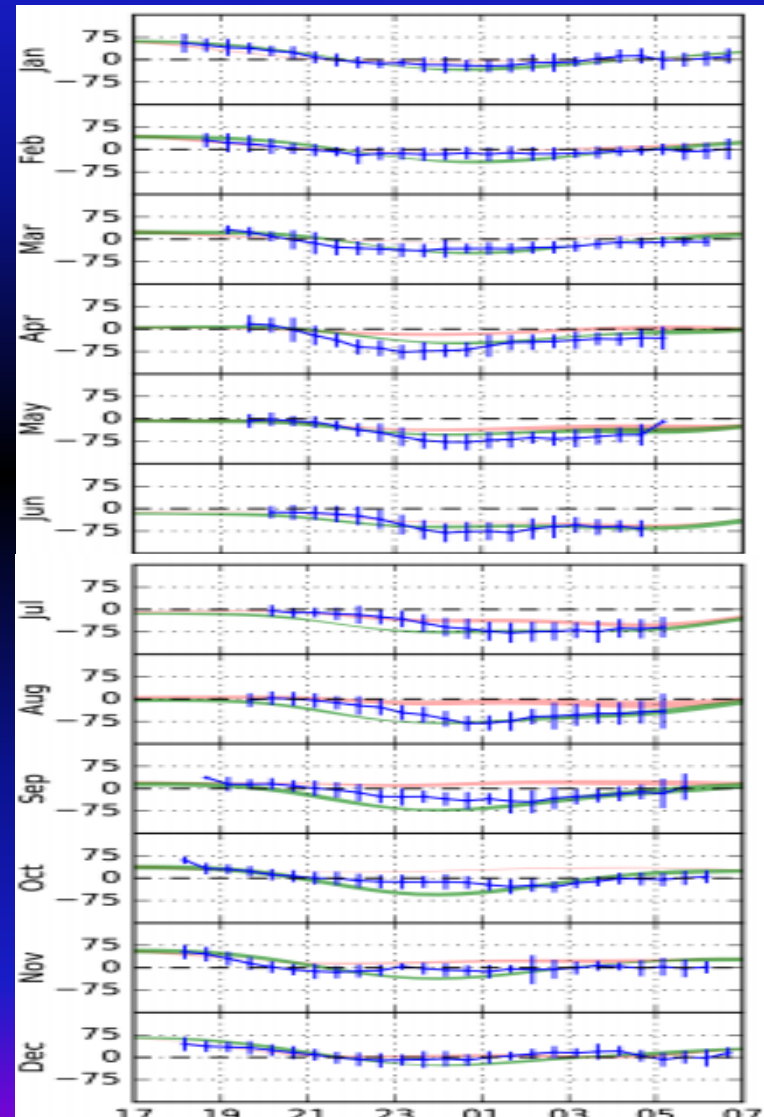
Atlas Mountains, Morocco 31.2°N, 7.9°E; geom: 22.8°N

Zonal Wind

Meridional Wind

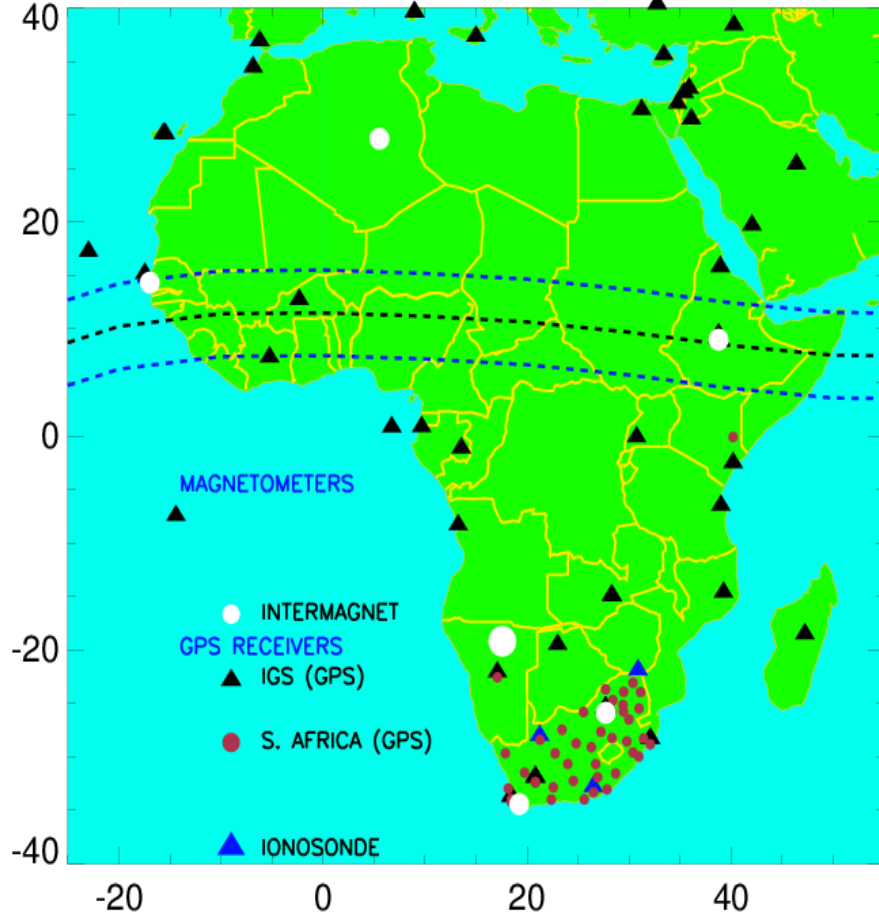


Data
HWM07
&
HWM14
models

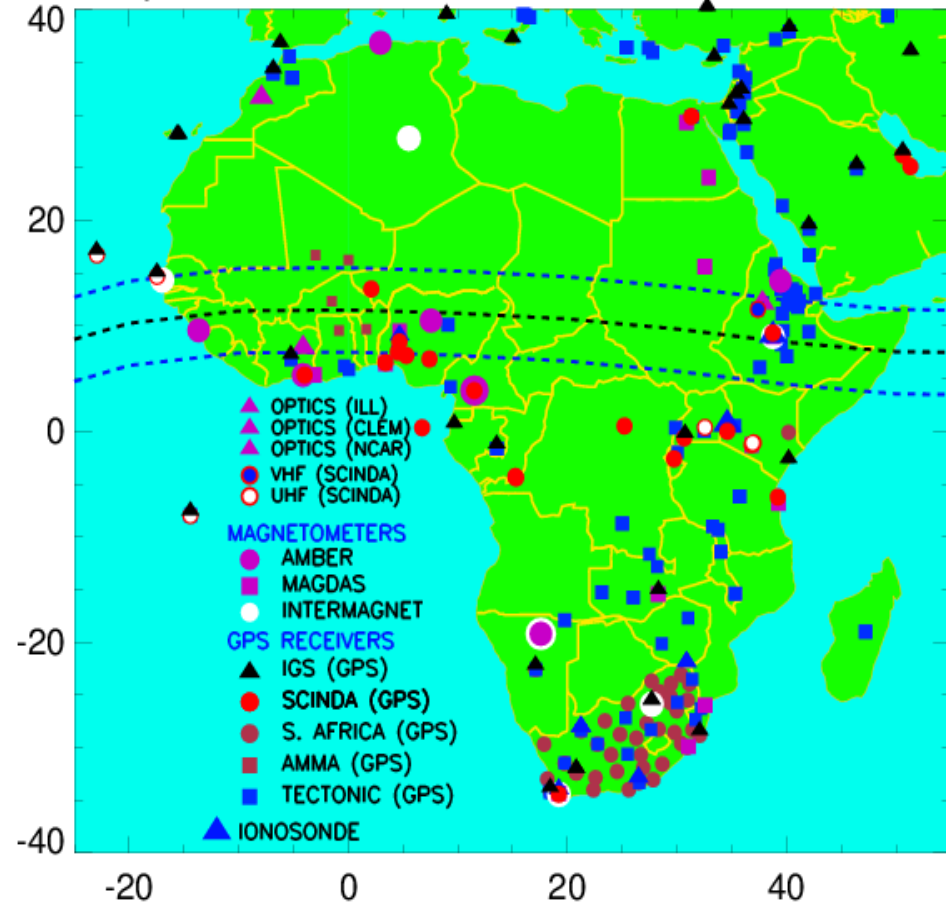


General Instrumentation in Africa

Space Science Instruments in Africa: 5 years ago



Space Science Instruments in Africa: Now



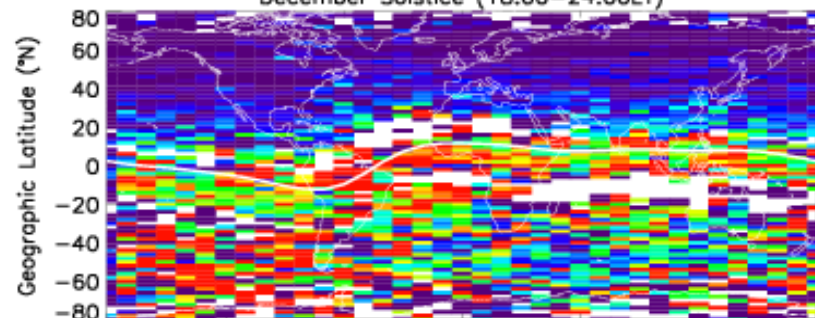
In 2007

In 2015

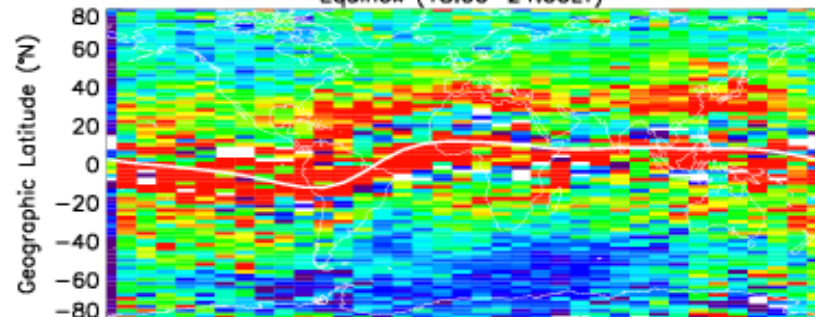
Thank You!

Irregularity Distributions from COSMIC RO data 2010–2014

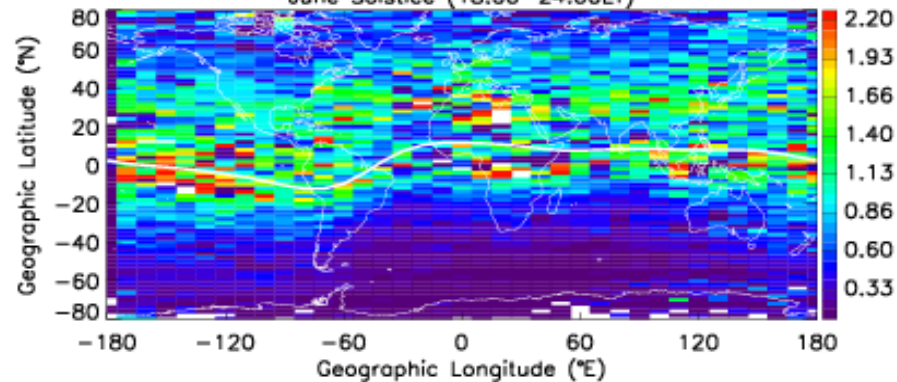
December Solstice (18:00–24:00LT)



Equinox (18:00–24:00LT)

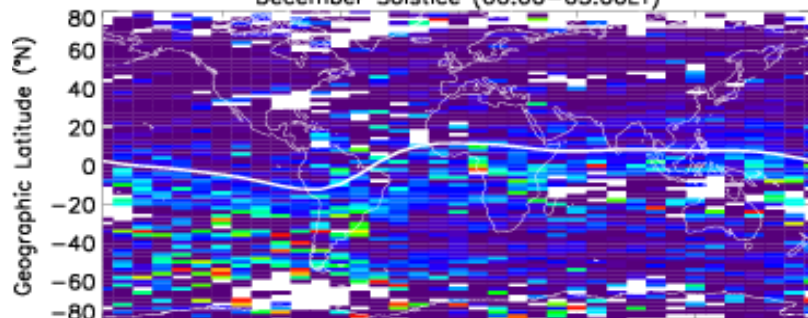


June Solstice (18:00–24:00LT)

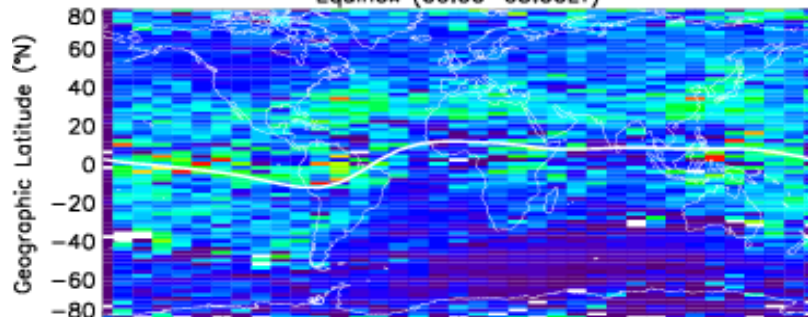


Irregularity Distributions from COSMIC RO data 2010–2014

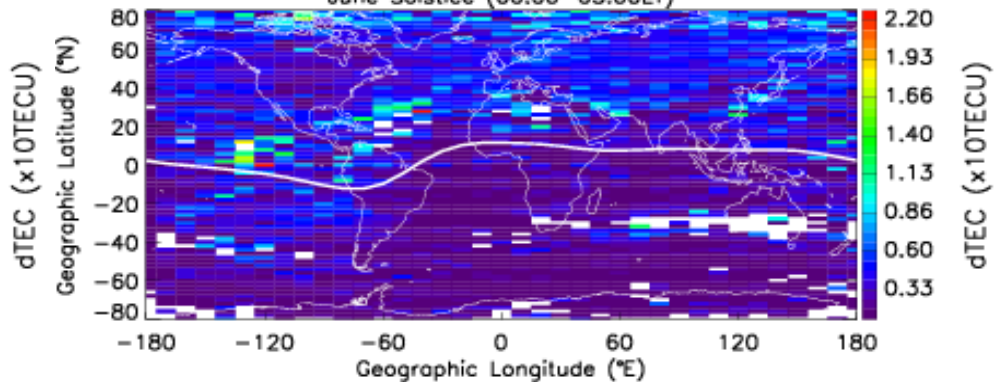
December Solstice (00:00–05:00LT)



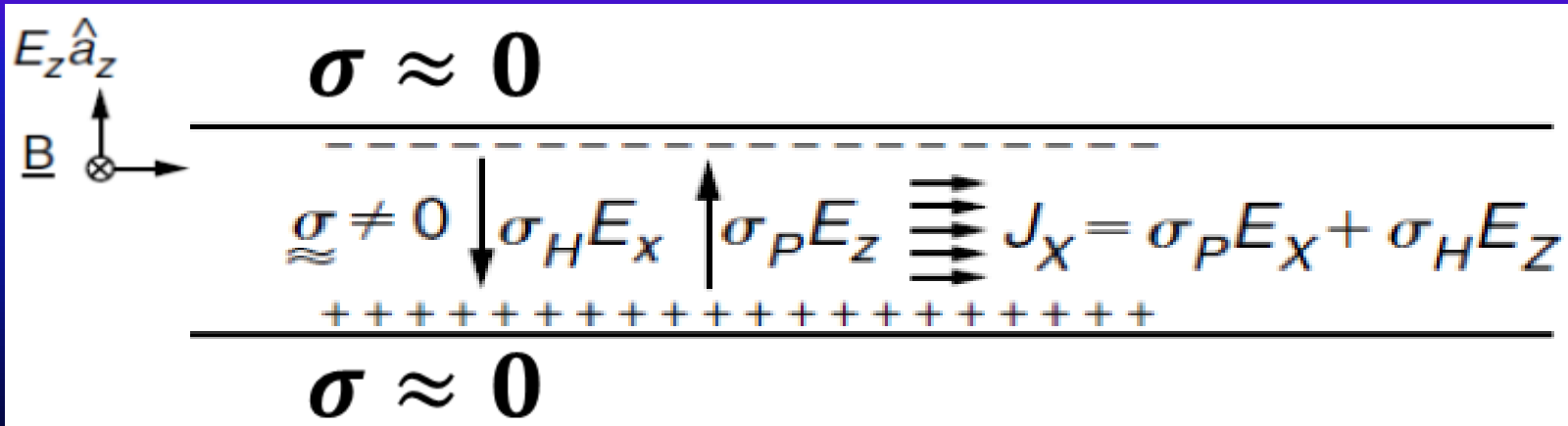
Equinox (00:00–05:00LT)



June Solstice (00:00–05:00LT)



Equatorial Electrojet (EEJ) ~ Equatorial Drift



$$J_z = -\sigma_H E_x + \sigma_P E_z = 0 \quad \longrightarrow \quad E_z = \frac{\sigma_H}{\sigma_P} E_x$$

$$J_x = \sigma_P E_x + \sigma_H \left(\frac{\sigma_H}{\sigma_P} E_x \right)$$

$$J_x = \left(\sigma_P + \frac{\sigma_H^2}{\sigma_P} \right) E_x = \sigma_C E_x \quad \longrightarrow \quad \text{EEJ current}$$

→ EEJ current, which is generated by Hall & Pederson conductivities in the vicinity of geomagnetic equator ($\pm 3^\circ$), is proportional to vertical drifts and can be estimated using magnetometer data.